

Third International Workshop  
NONLINEAR PROCESSES IN OCEANIC AND ATMOSPHERIC FLOWS

Evolution of secondary whirls in  
thermoconvective vortices in a route  
to chaos

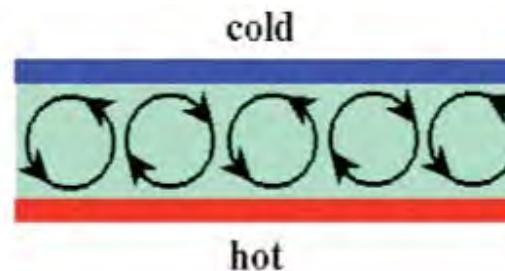
Henar Herrero

Dpto. Matemáticas, Universidad de Castilla-La Mancha, Spain  
(María Cruz Navarro, Damián Castaño UCLM)

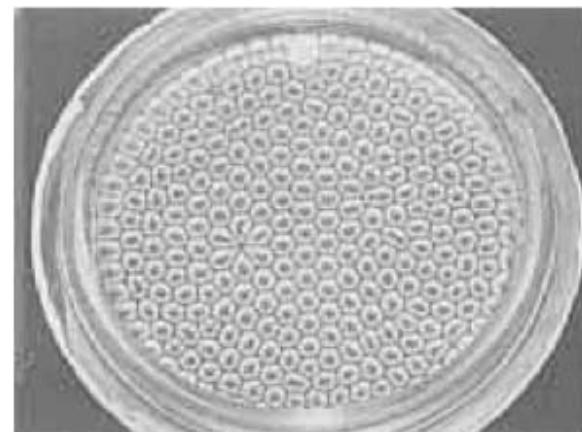


## Background

### Rayleigh-Bénard convection



*H. Bénard, 1900*

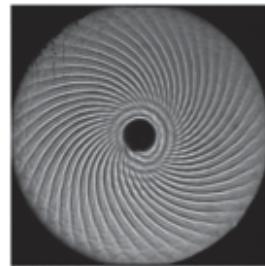


## Background

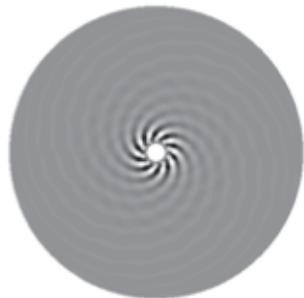
Localized heating in a cylindrical annulus and small cell



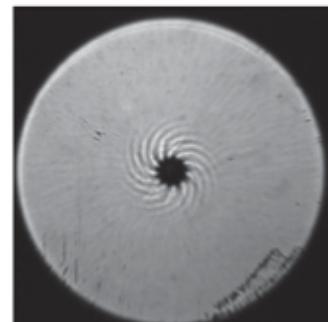
S. Hoyas, H. Herrero y A. M.  
Mancho , Phys. Fluids (2005)



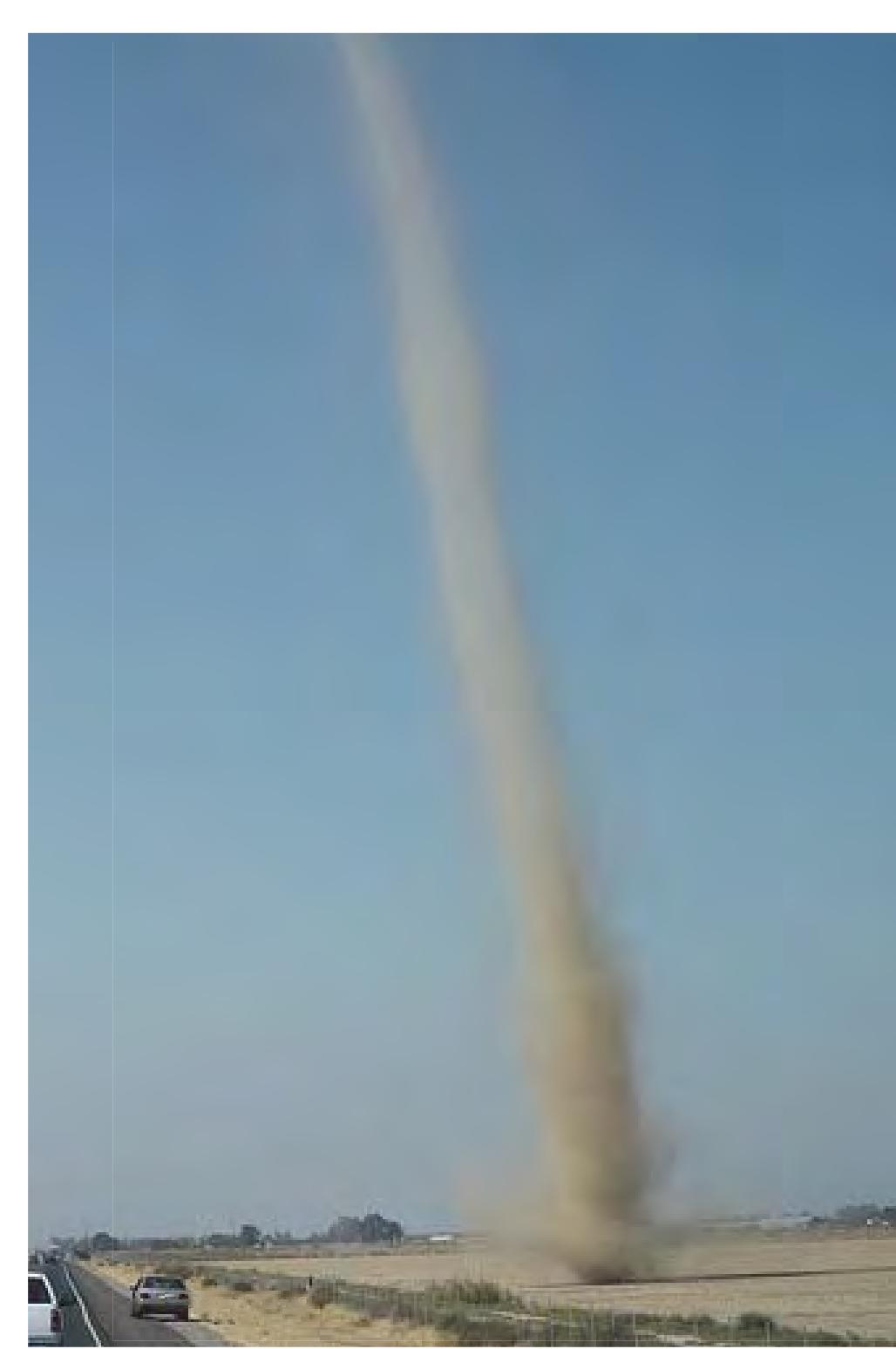
N. Garnier y A. Chiffaudel, Eur.  
Phys. J. (2001)



S. Hoyas, et al.  
Phys. of Fluids (2005)



Waves,  
mass transport?



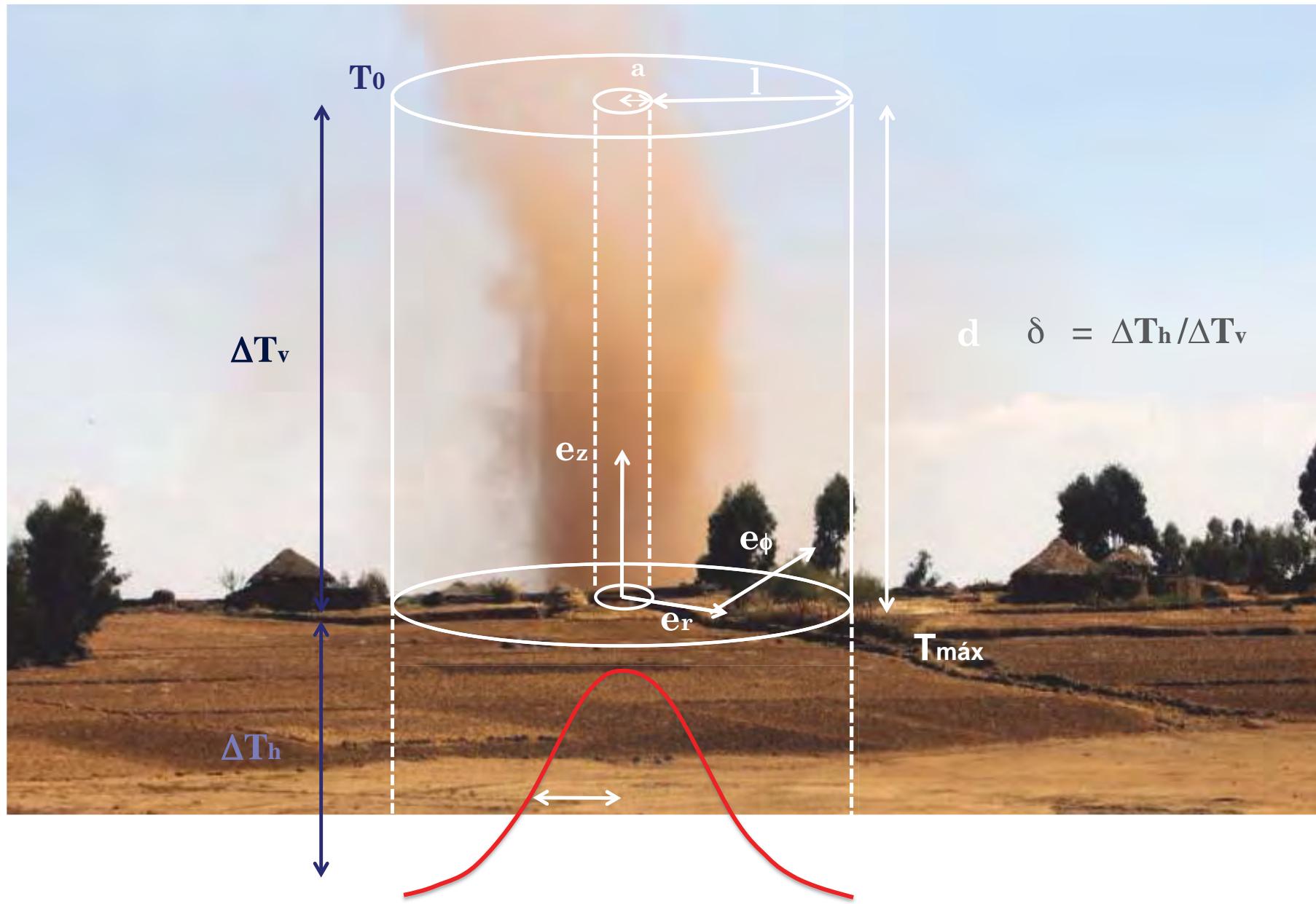
- Localized heating in a cylindrical domain,
- we look for mass transport,
- approximation to atmospheric phenomena

## **Dust devil**

# Outline

- Formulation
- Numerical methods
- Numerical results
- Comparison with observations

# Domain



## Hidrodynamic equations + Heat equation

$$\nabla \cdot \mathbf{u} = 0$$

$$\partial_t \Theta + \mathbf{u} \cdot \nabla \Theta = \nabla^2 \Theta$$

$$\partial_t \mathbf{u} + (\mathbf{u} \cdot \nabla) \mathbf{u} = \cancel{Pr} (-\nabla p + \nabla^2 \mathbf{u} + \cancel{R} \Theta \mathbf{e}_z)$$

0.7 air

# Cylindrical annulus

Inner cylinder

No slip

$$u_r = 0$$

$$u_\phi = 0$$

$$u_z = 0$$

Insulating wall

$$\partial_r \Theta = 0$$

$$\mathbf{u} = 0$$

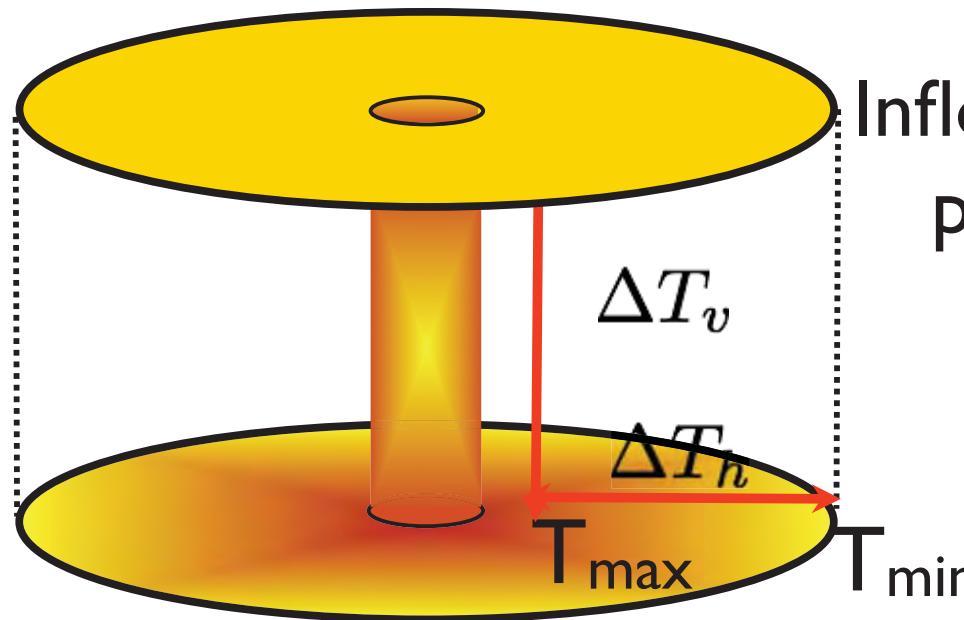
$$T_0$$

Inflow/outflow  
permitted

$$\partial_r u_r = 0$$

$$\partial_r u_\phi = 0$$

$$\partial_r u_z = 0$$



$$\partial_z u_r = \partial_z u_\phi = u_z = 0$$

Insulating

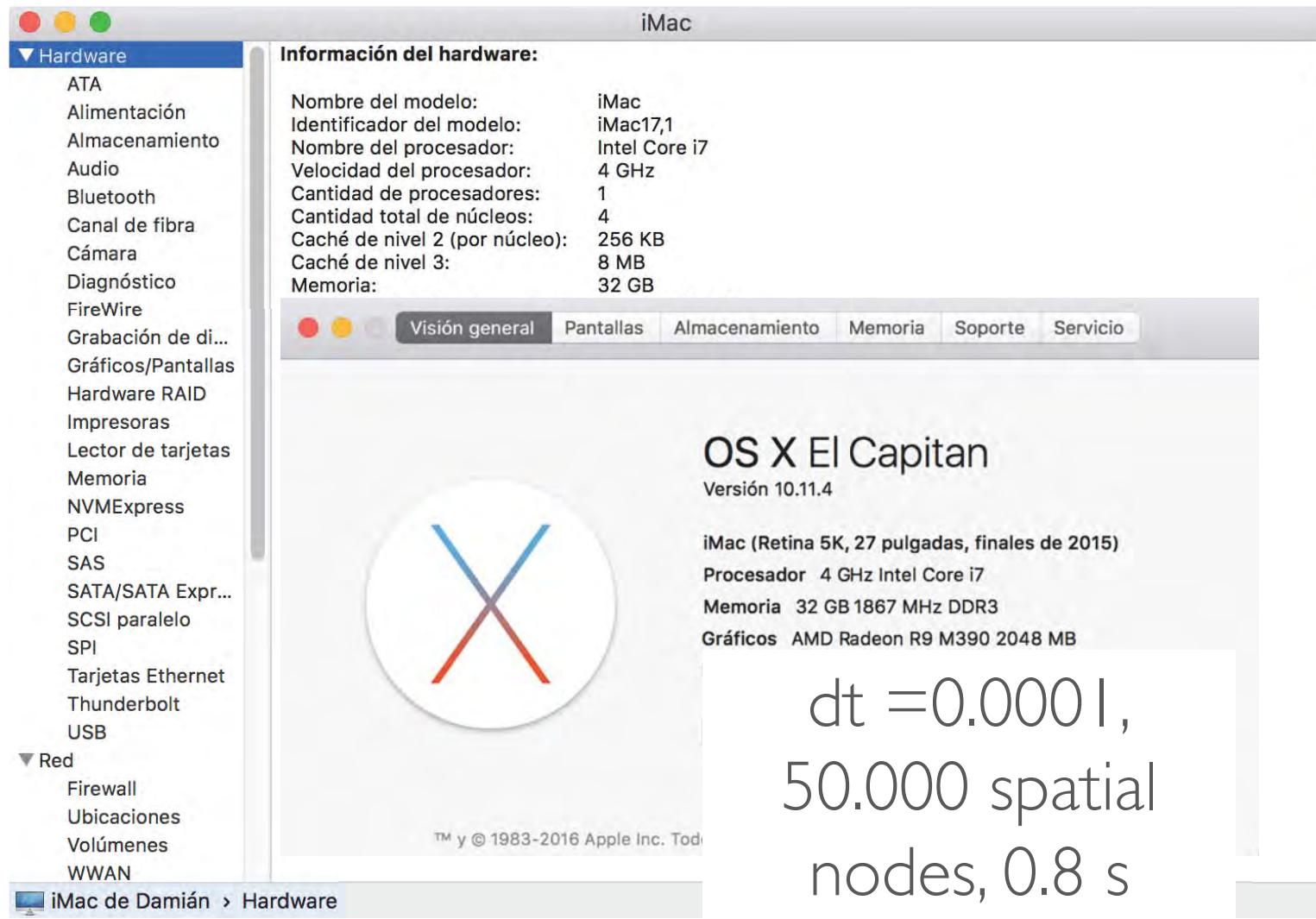
$$\partial_r \Theta = 0$$

$$\Theta = 1 - \delta(e^{(\frac{1}{\beta})^2} - e^{(\frac{1}{\beta} - (\frac{r-\bar{a}}{\Gamma})^2 \frac{1}{\beta})^2})/(e^{(\frac{1}{\beta})^2} - 1)$$

# Time dependent problem

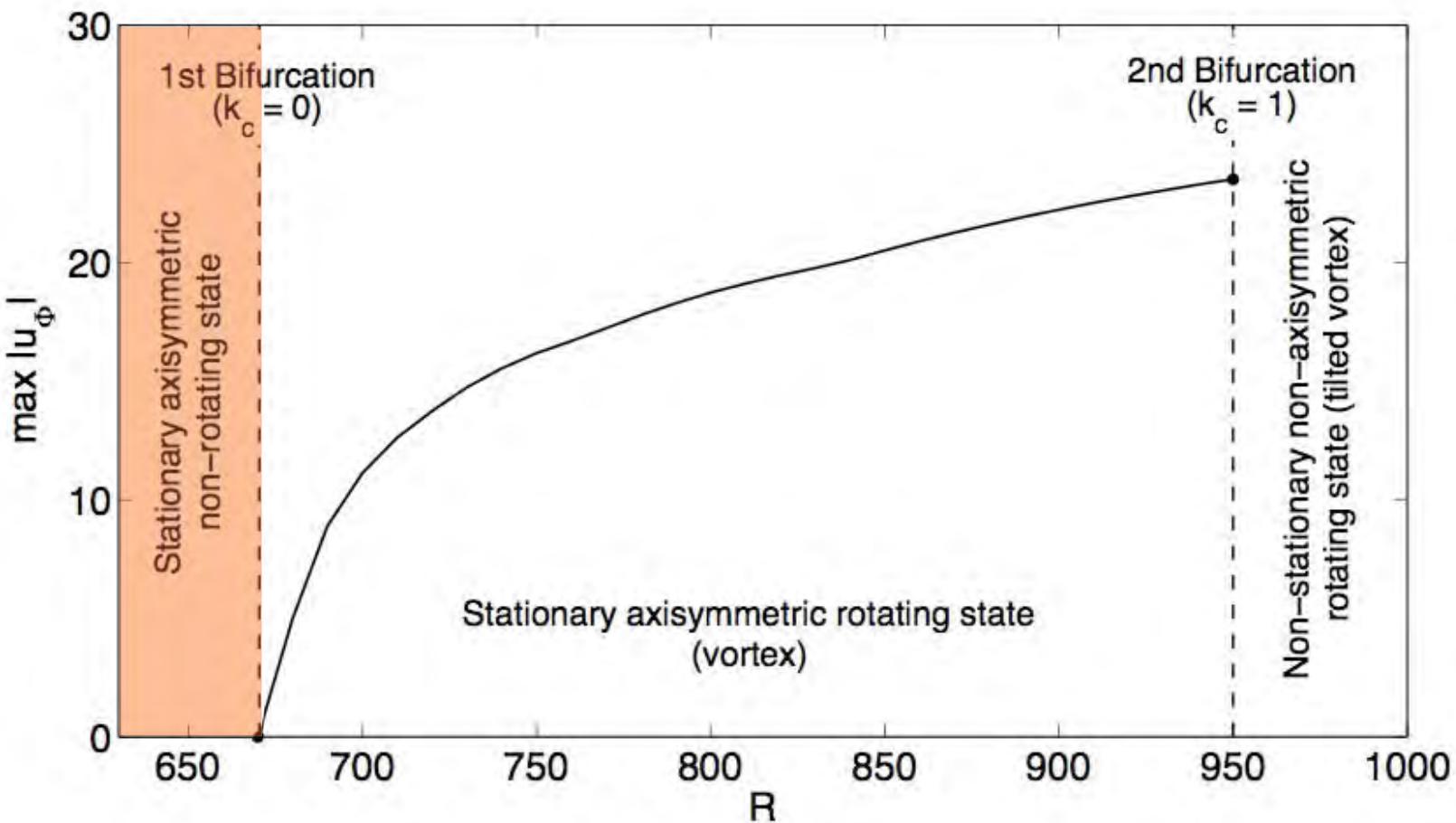
- Combination of Adams-Bashforth and backward differentiation formula (AB/BDF) scheme for time evolution (second order)
- Fourier and Chebyshev collocation
- Predictor-corrector scheme
  - Predictor Poisson for pressure
  - Predictor Helmholtz por velocity
  - Corrector

# 3. Time dependent problem

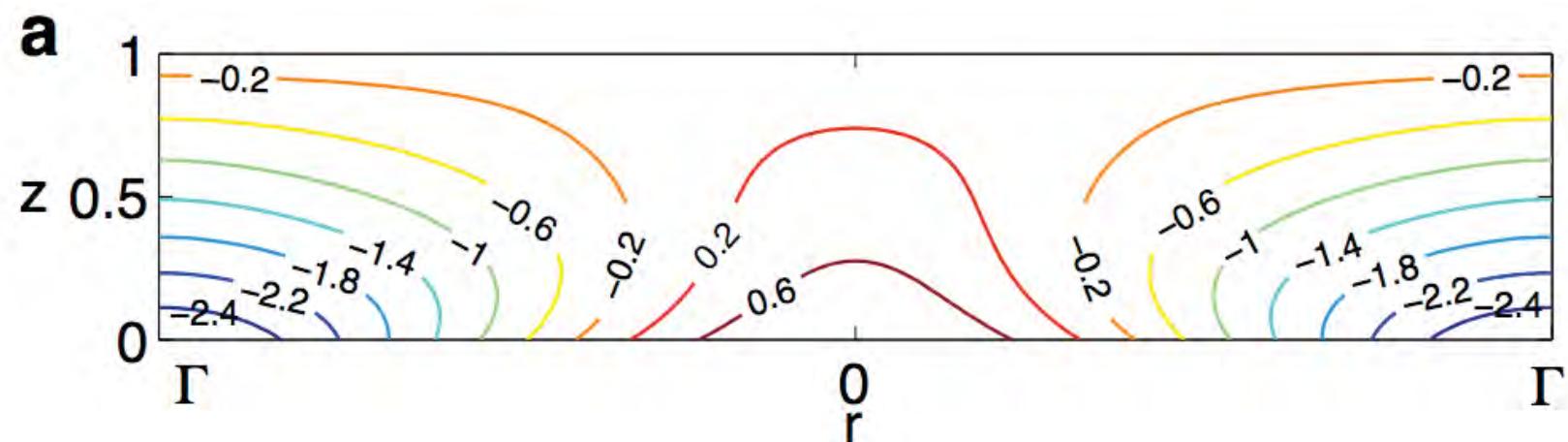


Plume

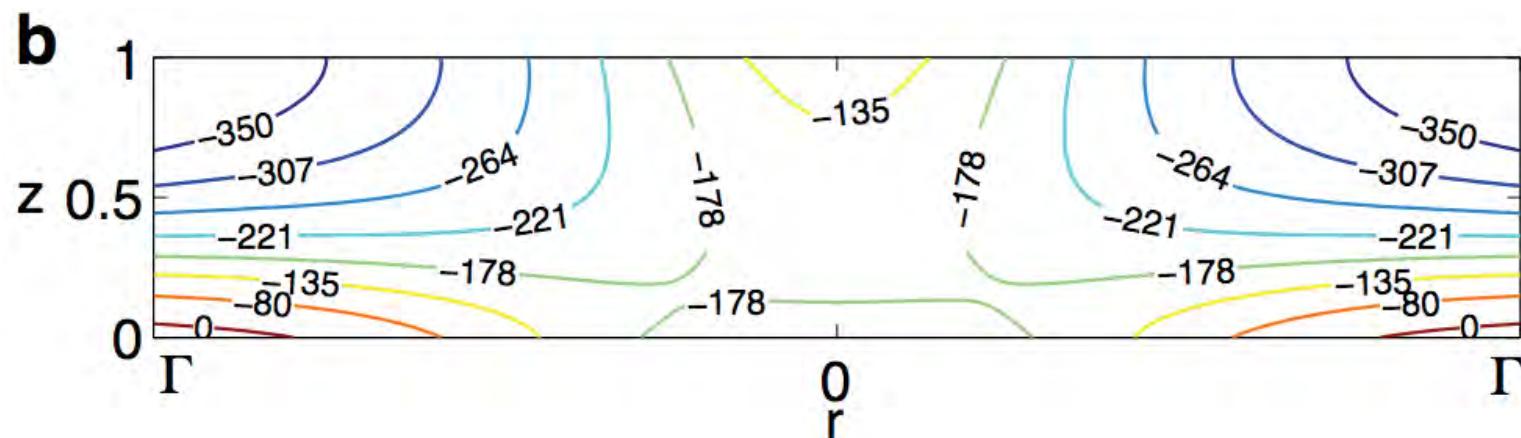
# Bifurcation diagram



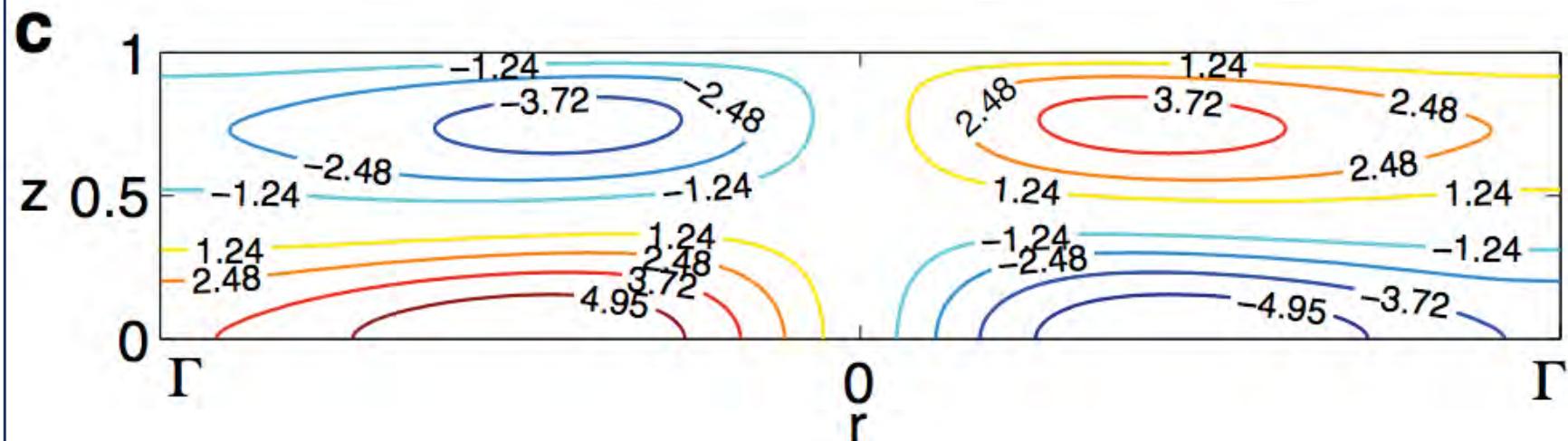
## Temperature



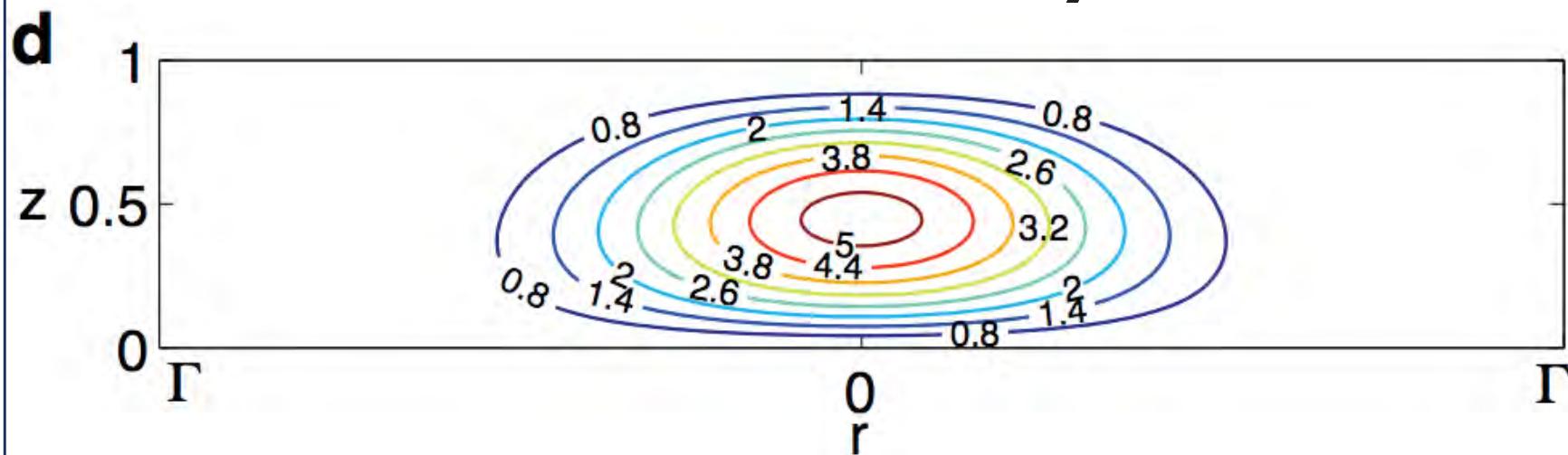
## Pressure



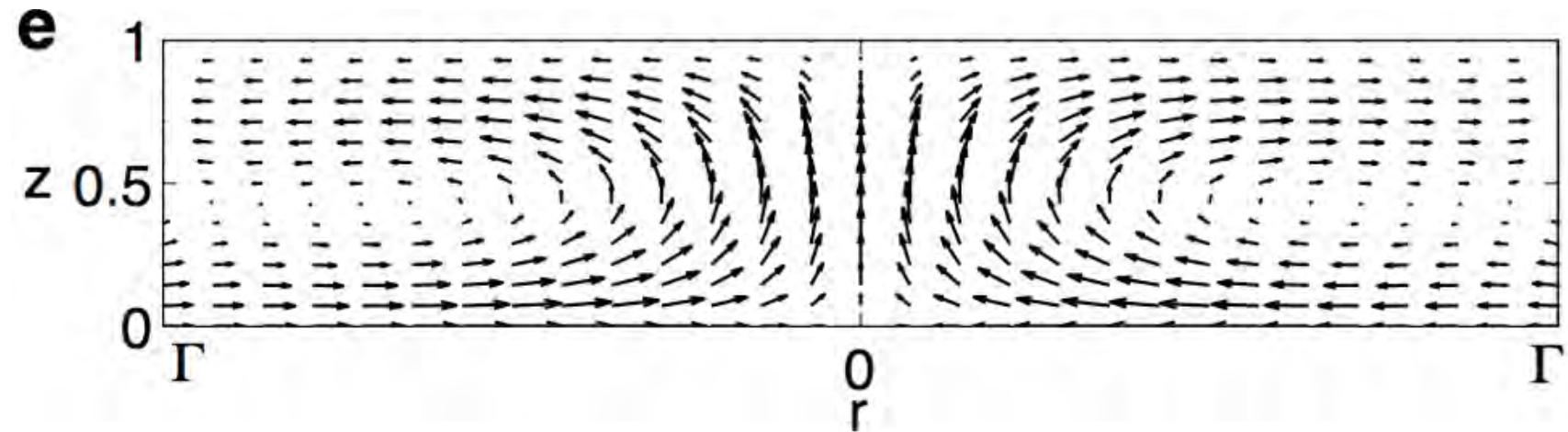
## Radial velocity



## Vertical velocity

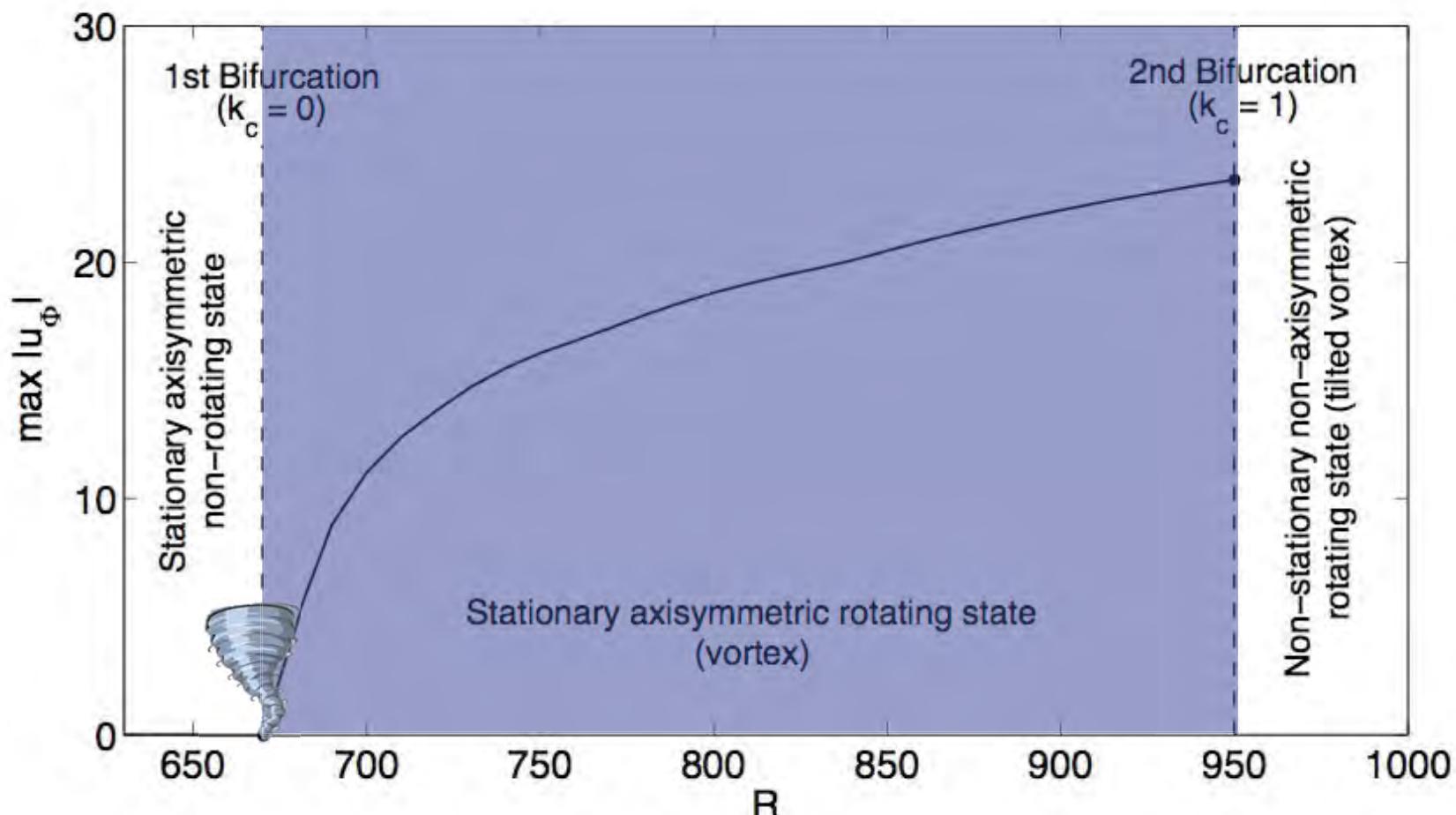


## **Velocity field**



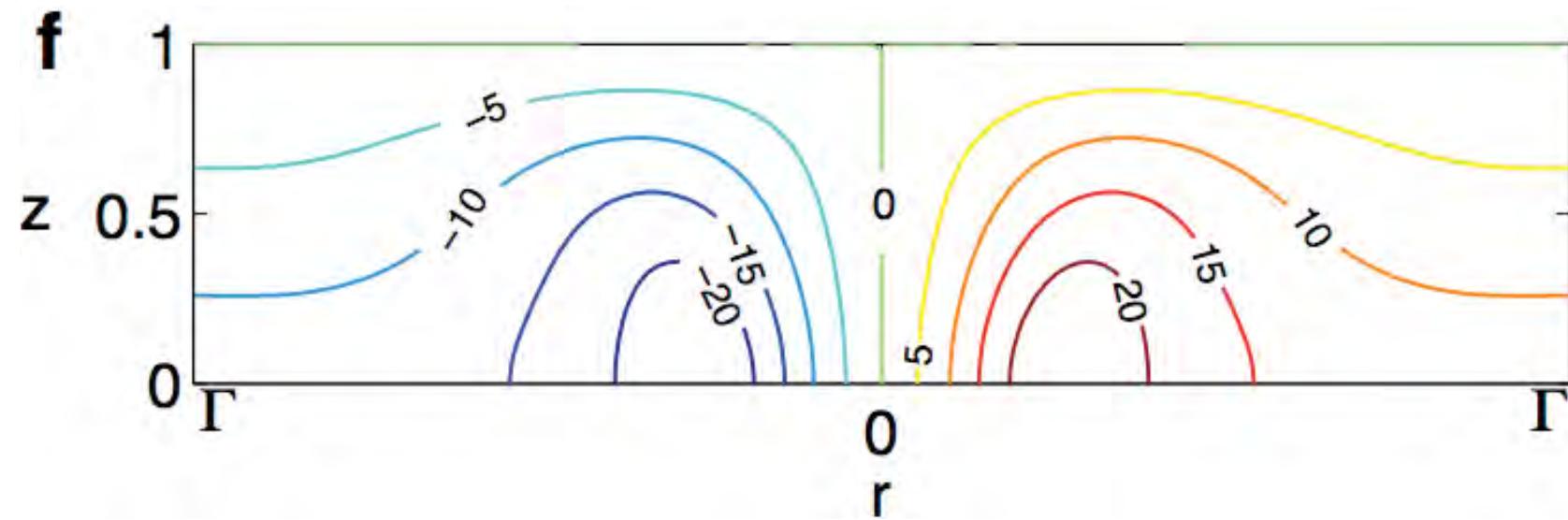
**NO azimuthal velocity**

# Bifurcation diagram



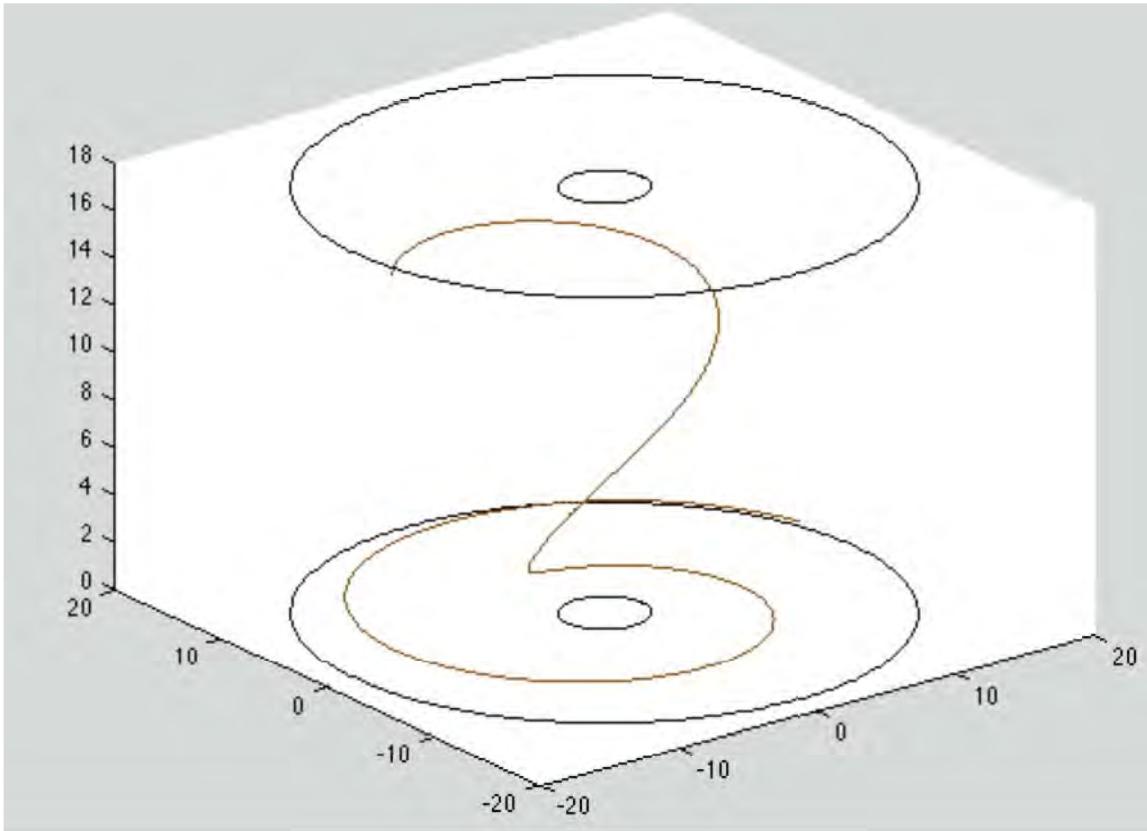
Vertical vortex

# Azimuthal velocity



Spin motion

# Trajectory of particles



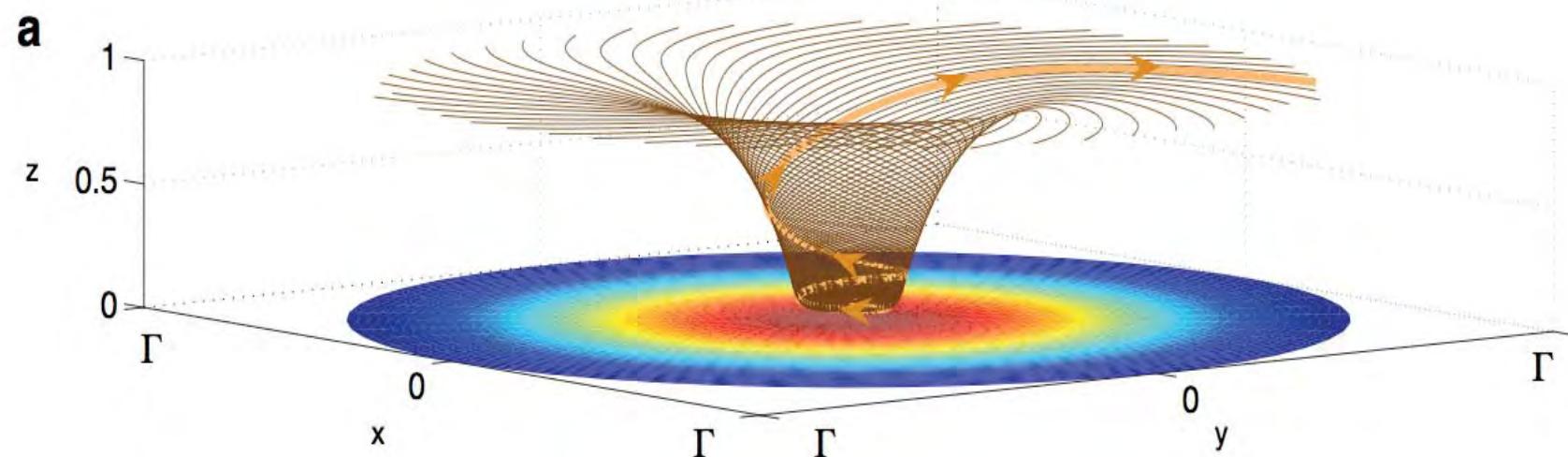
$$\frac{\partial r}{\partial t} = u_r(r, z)$$

$$\frac{\partial \phi}{\partial t} = u_\phi(r, z)$$

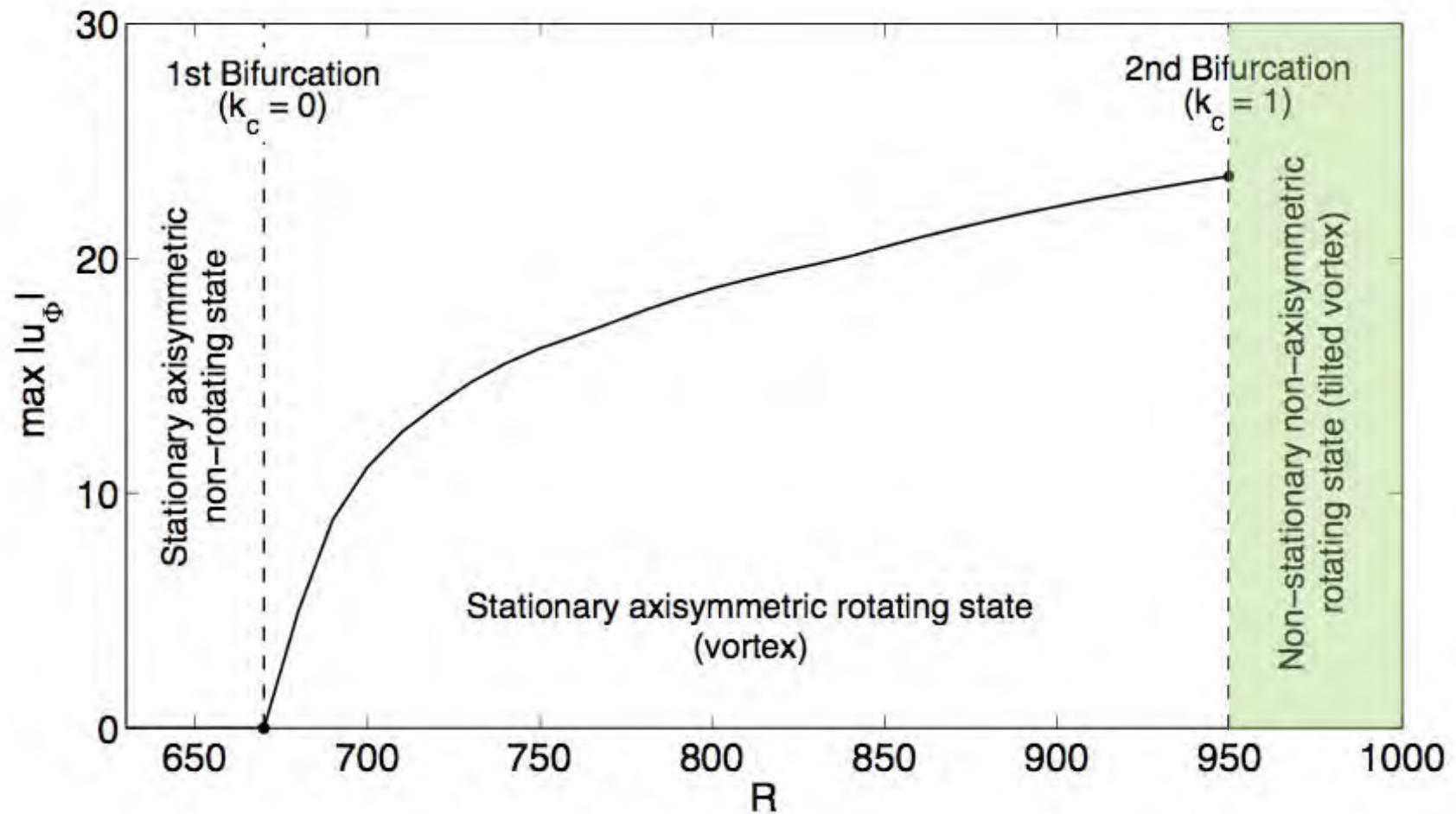
$$\frac{\partial z}{\partial t} = u_z(r, z)$$

Integrating the velocity fields with respect to time

# Trajectory of particles

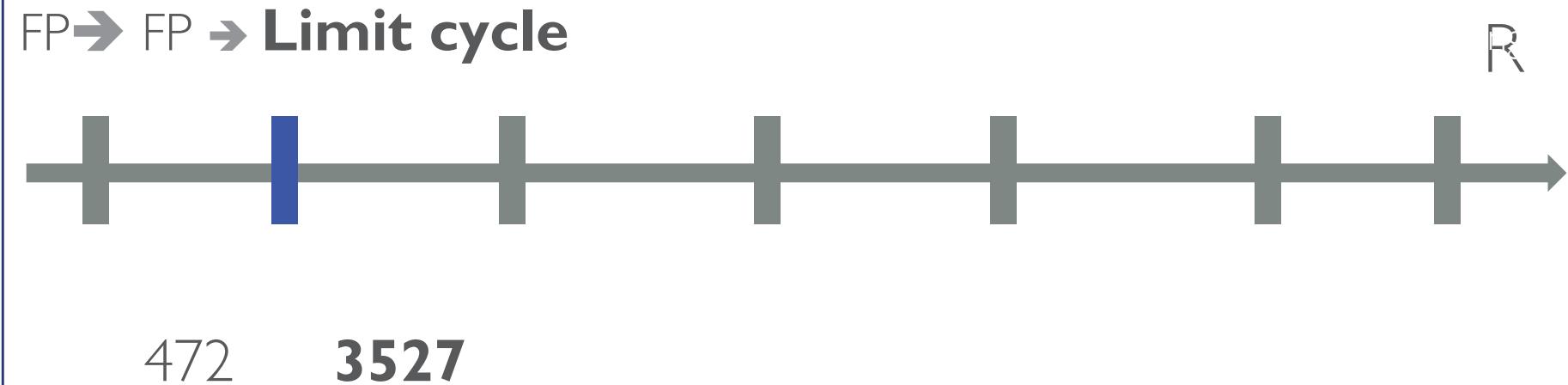


# Oscillatory Bifurcation diagram

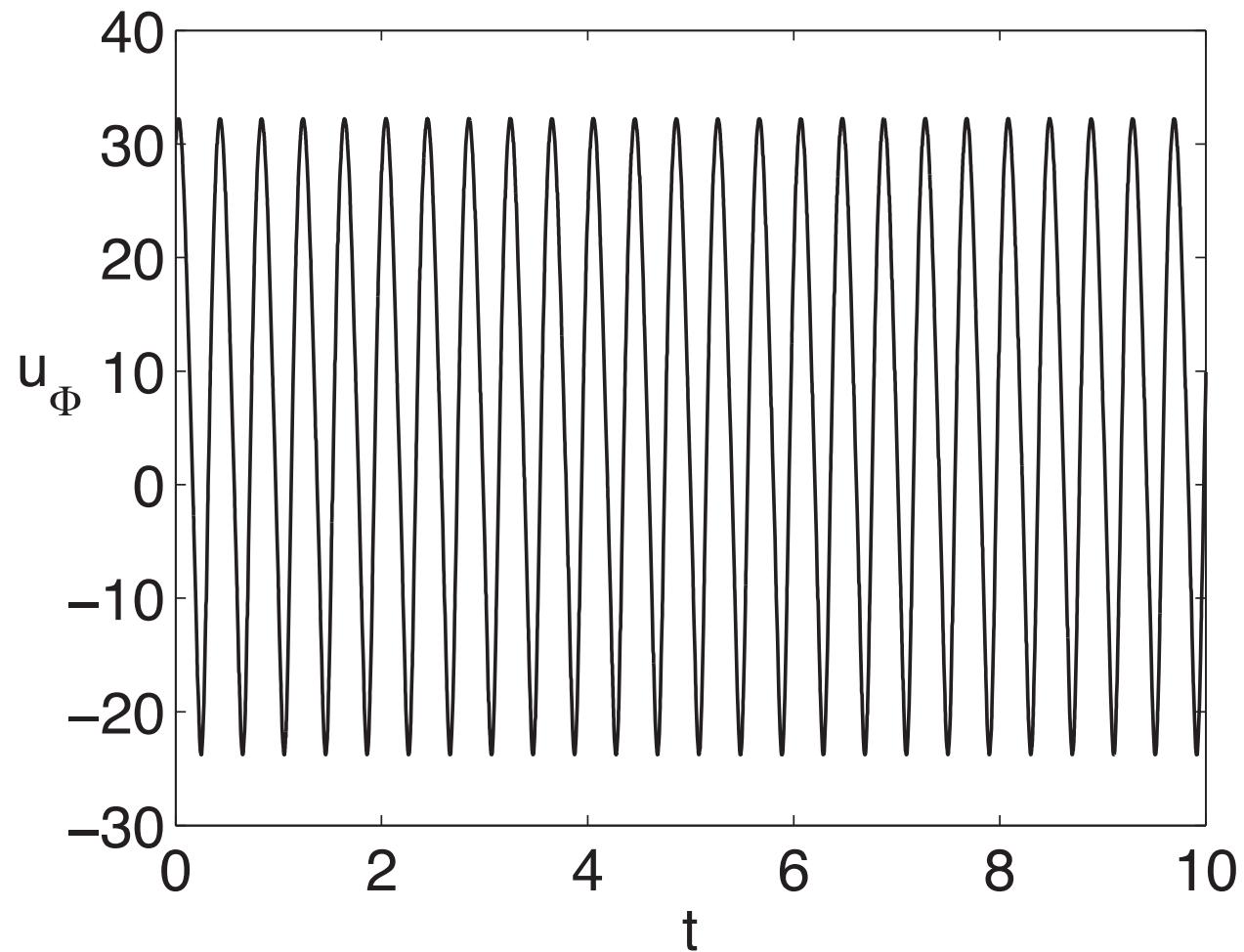


# Bifurcation diagram scheme

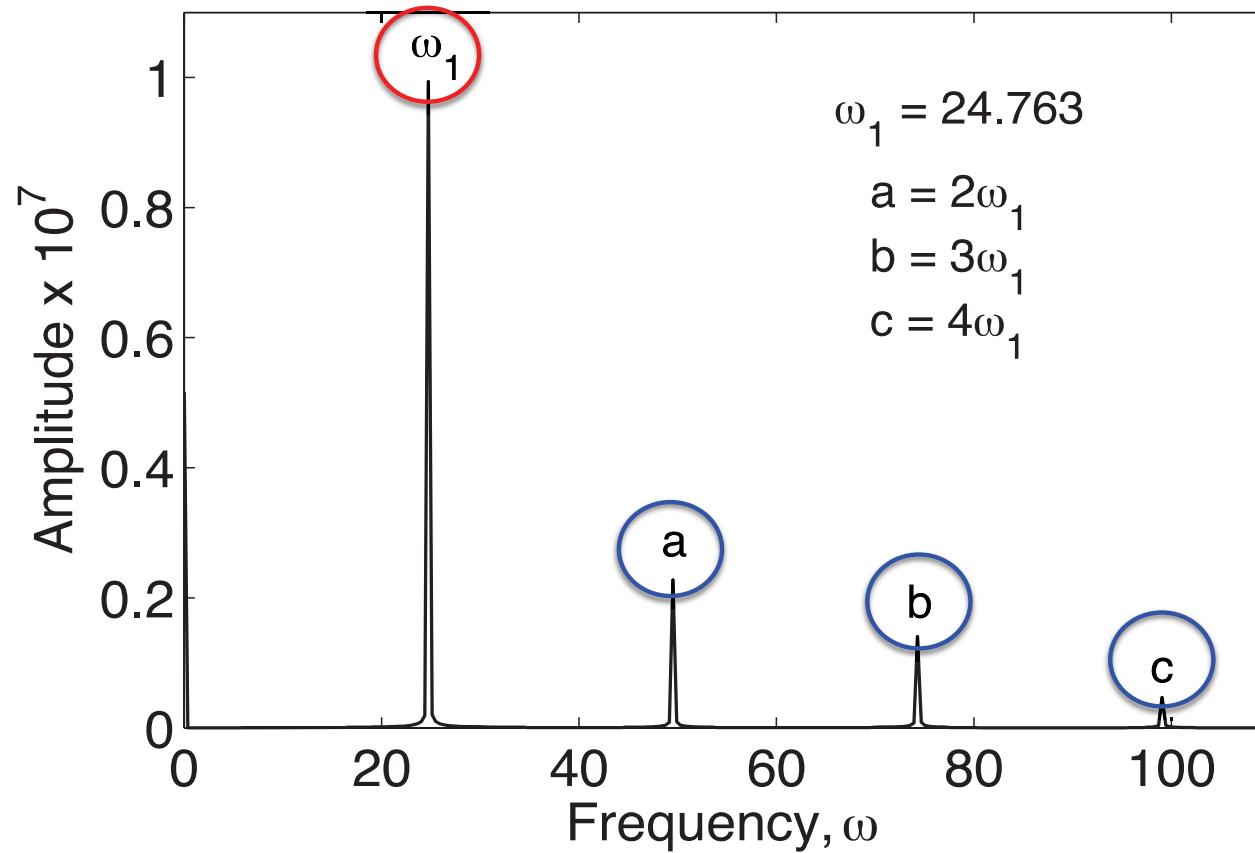
FP → FP → Limit cycle



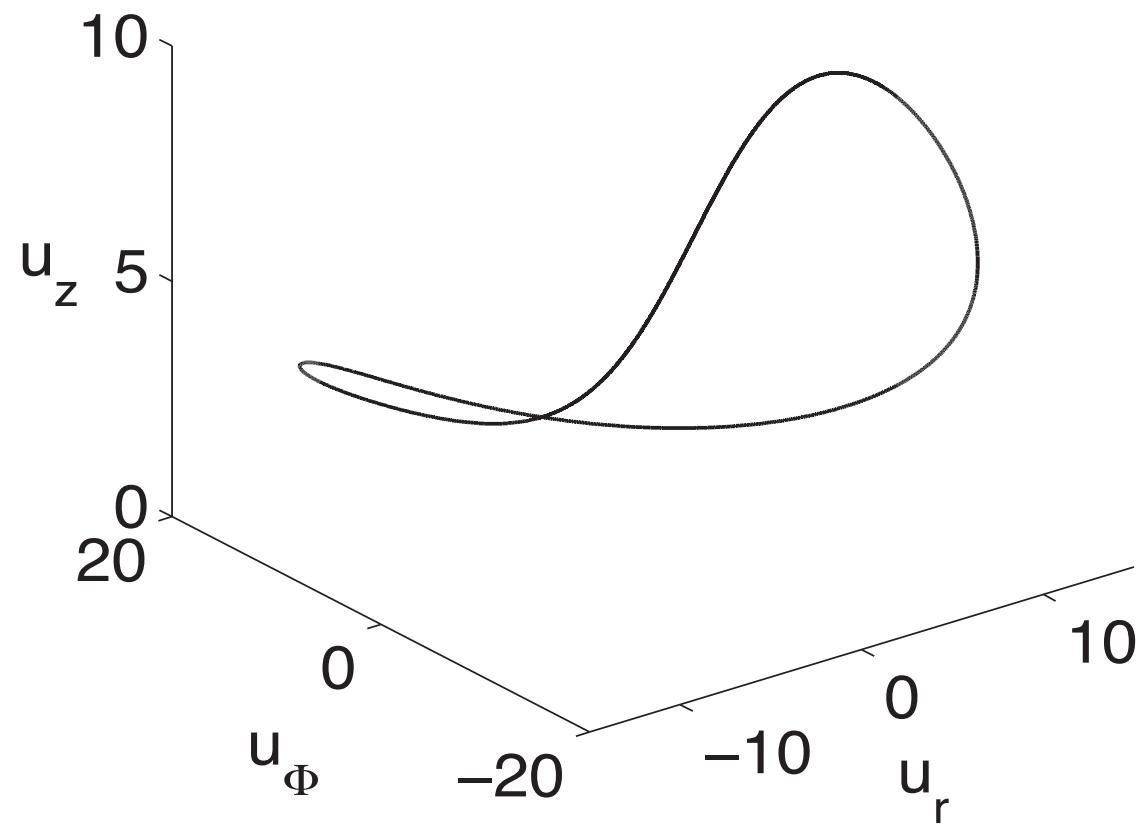
$\text{Ra} = 10000$

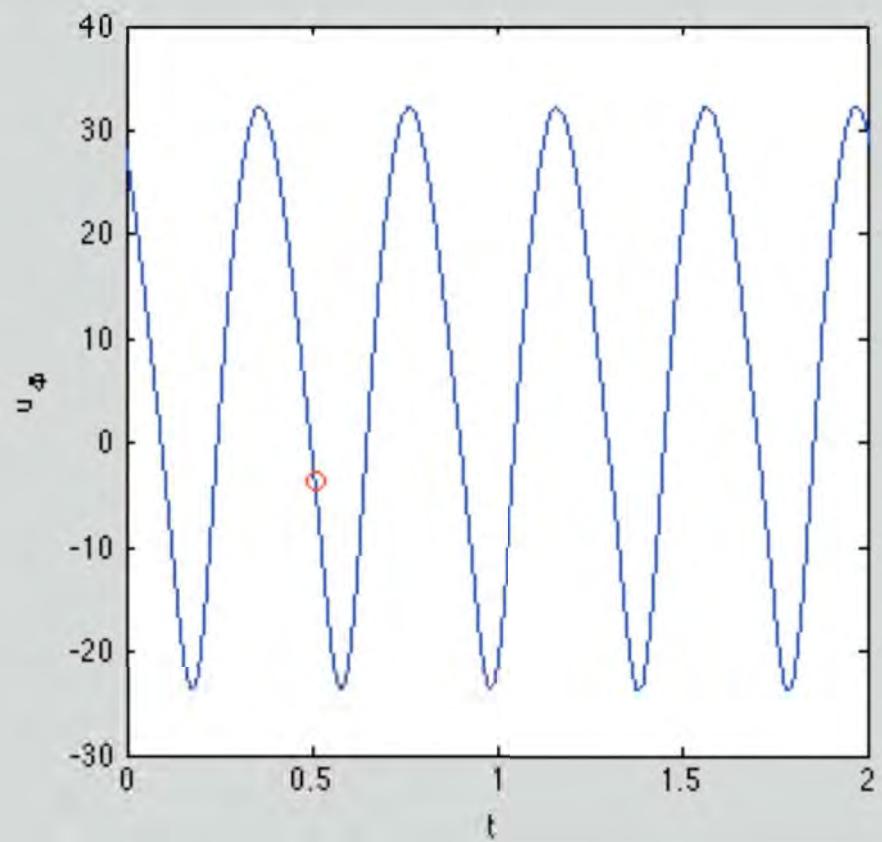
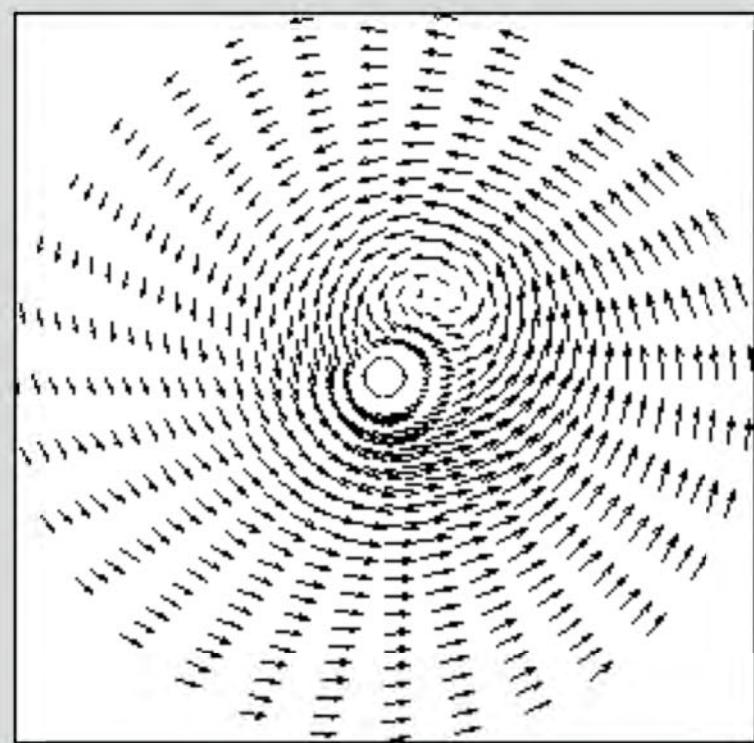


$\text{Ra} = 10000$

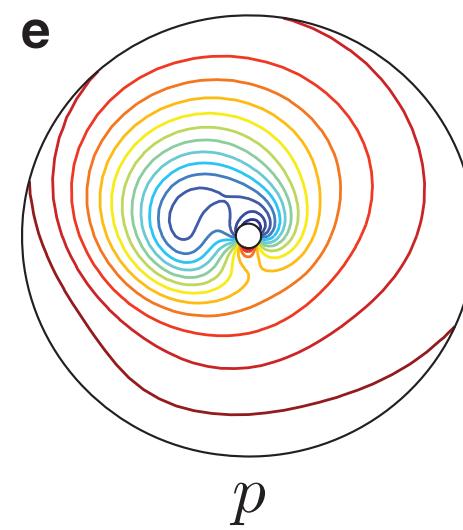
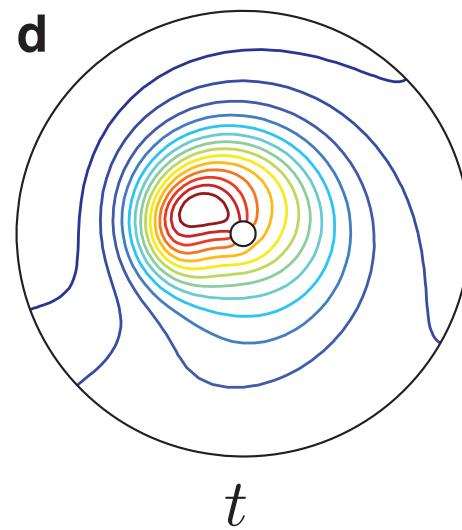
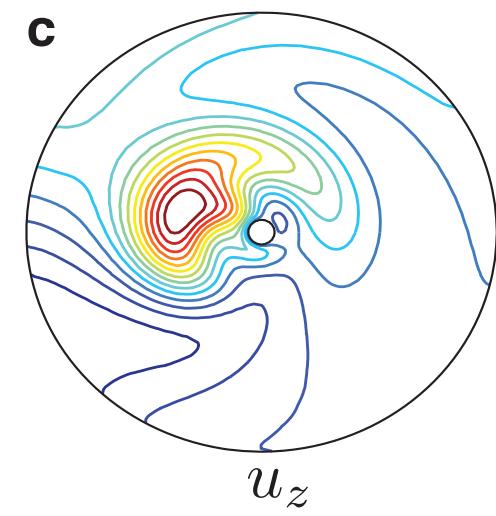
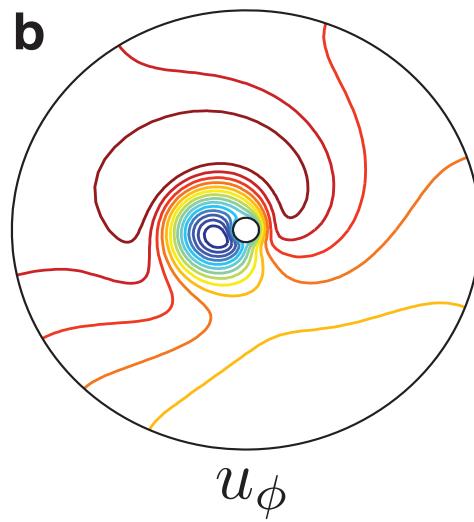
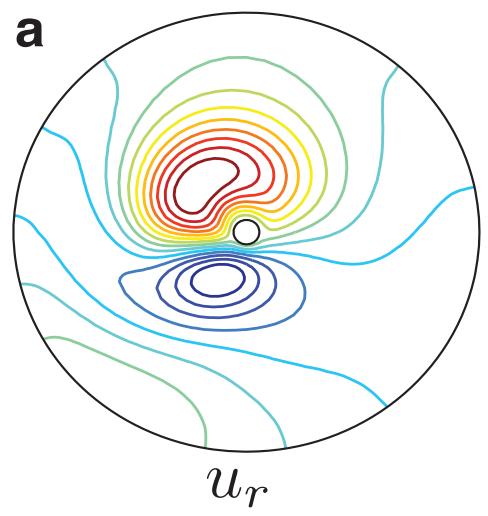


$\text{Ra} = 10000$

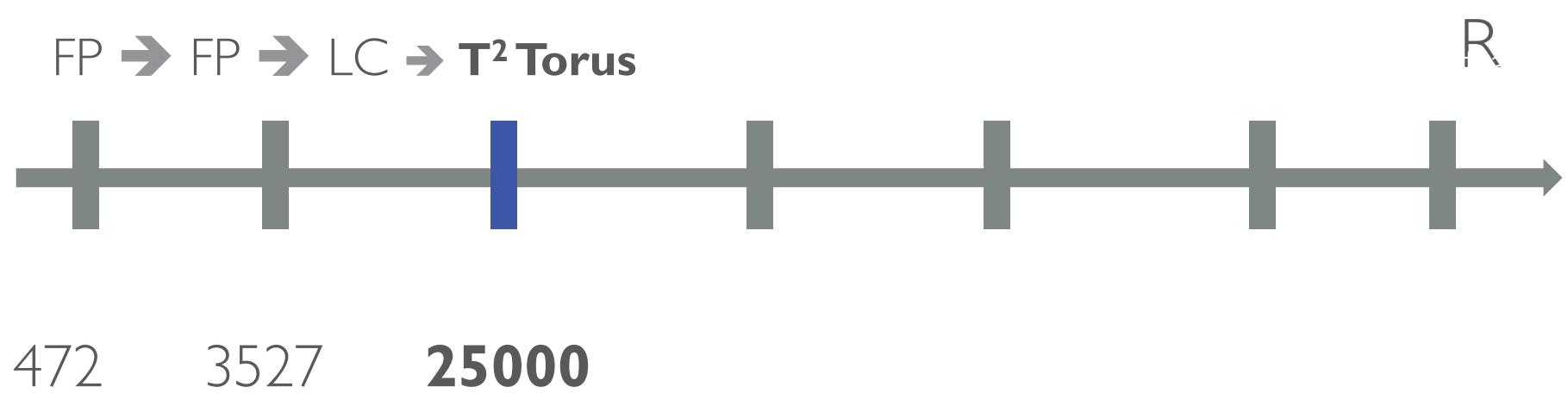




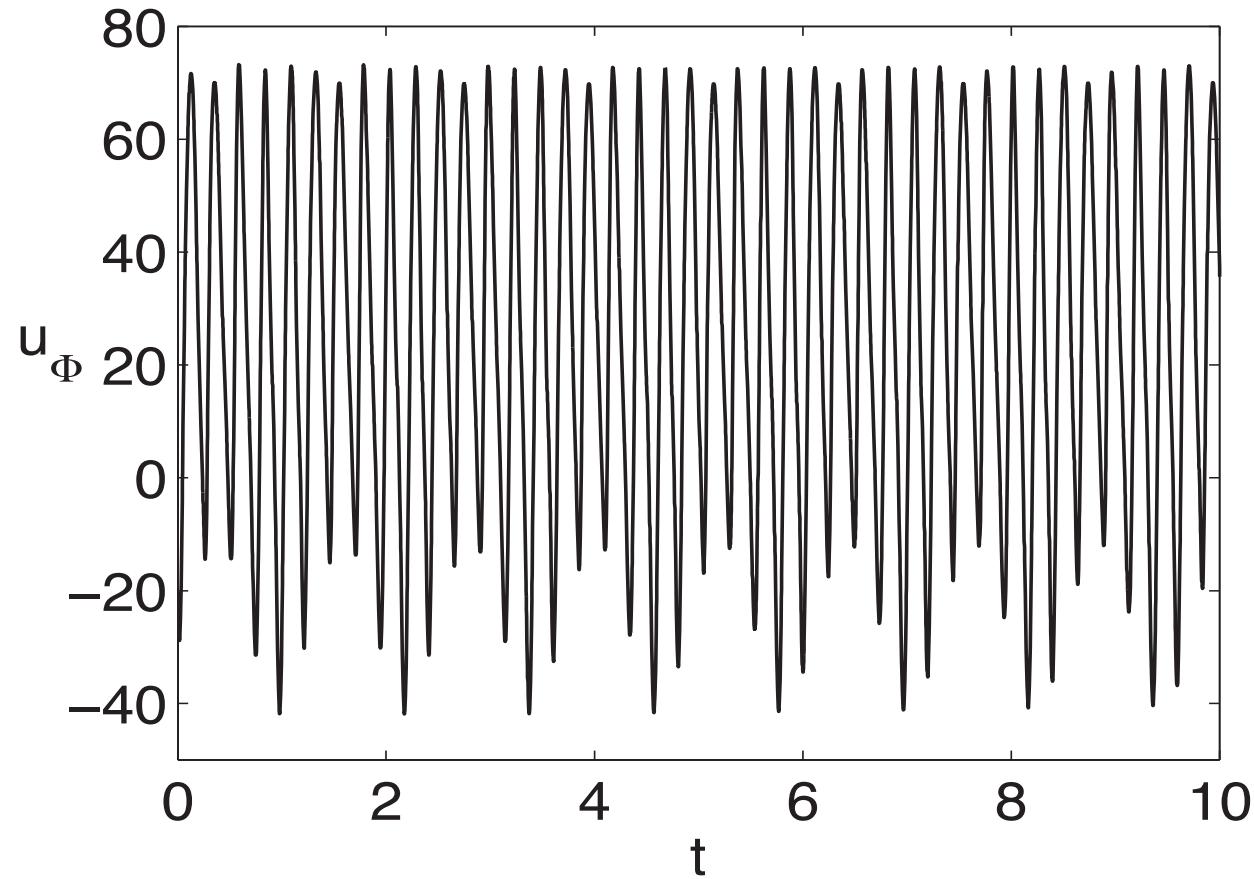
$\text{Ra} = 10000$



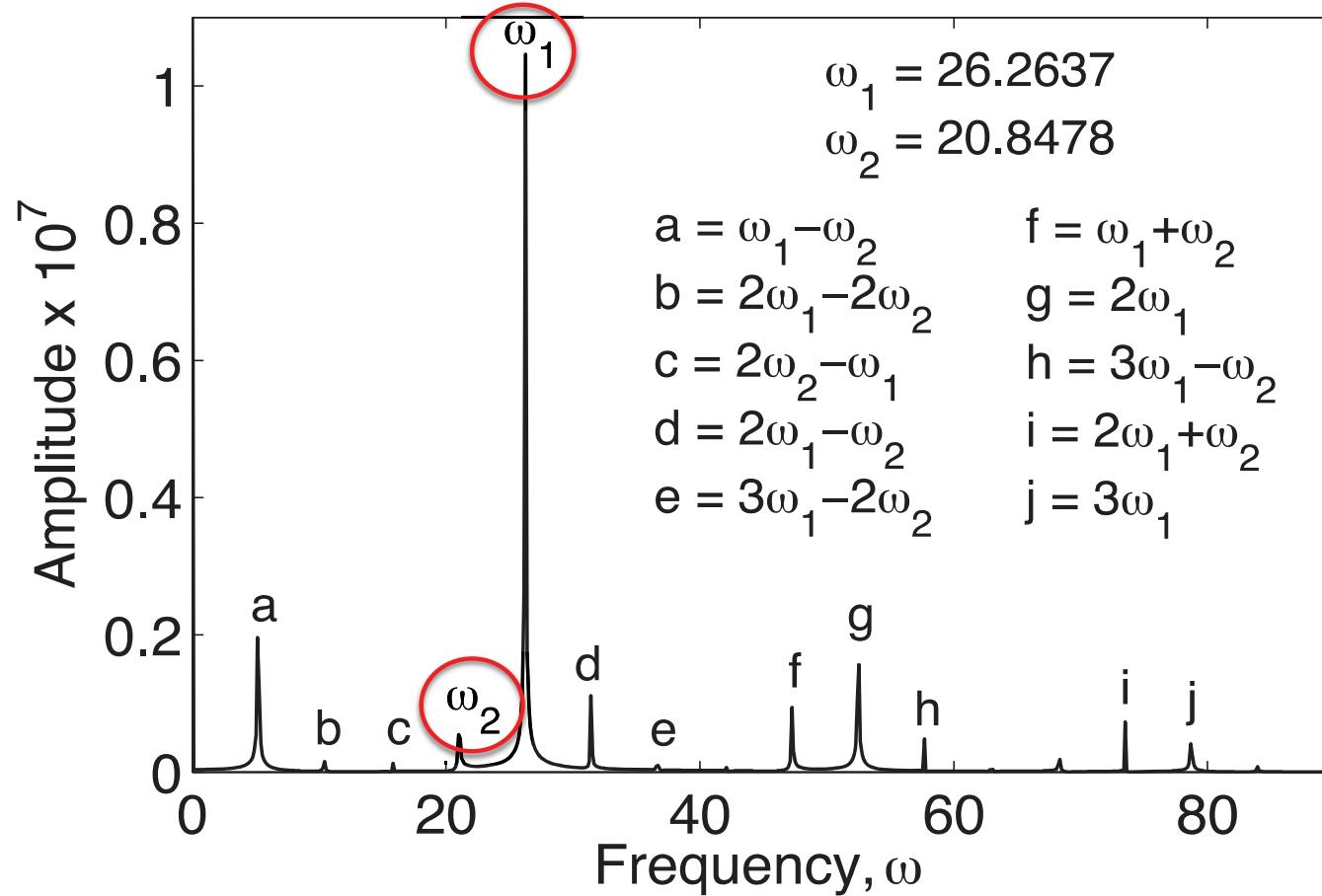
# Bifurcation diagram scheme



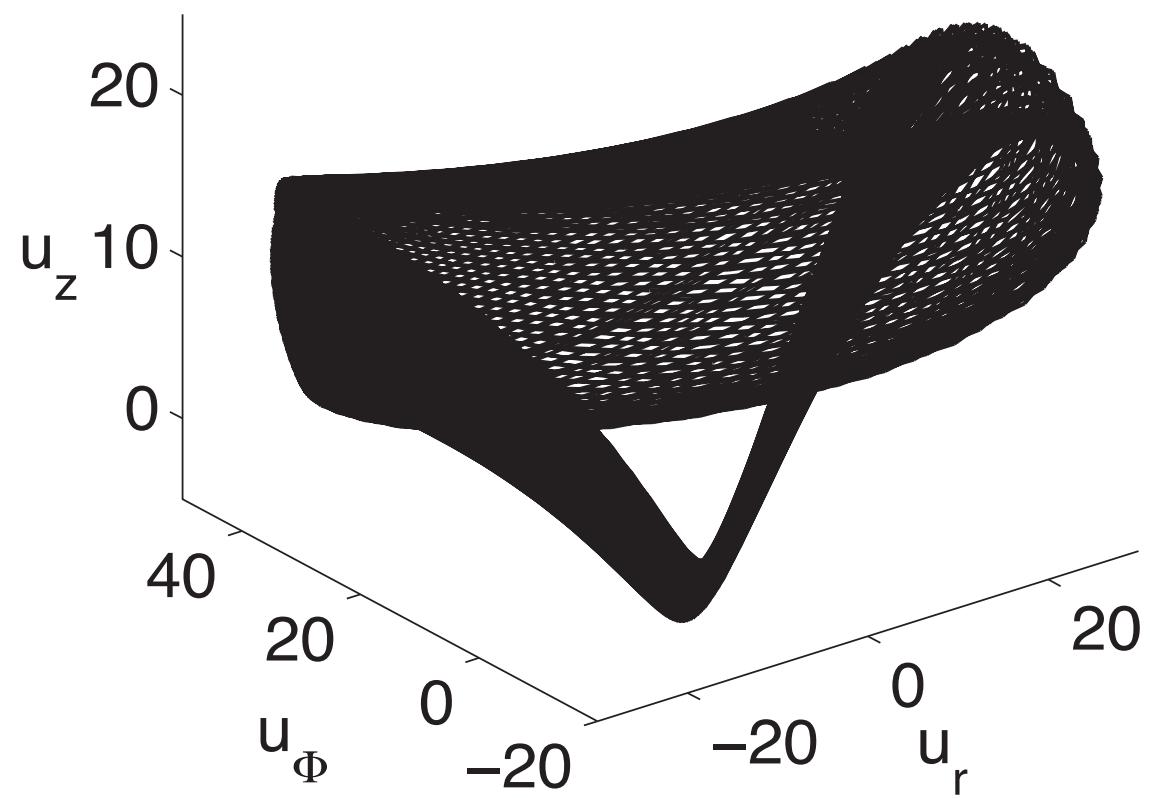
$\text{Ra} = 35000$

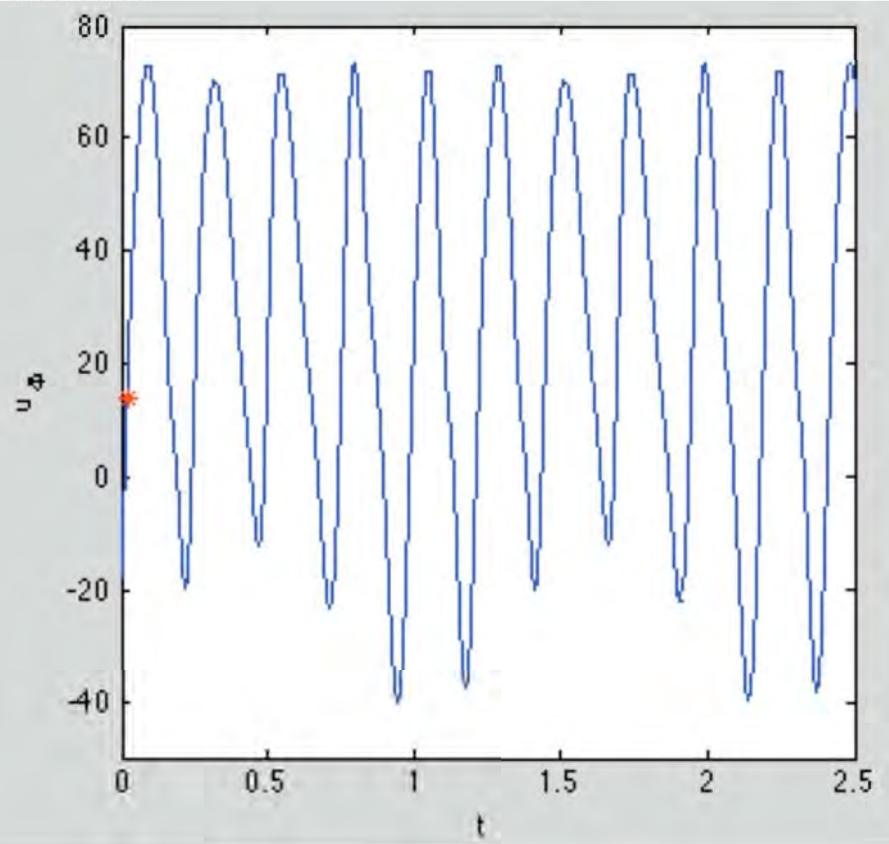
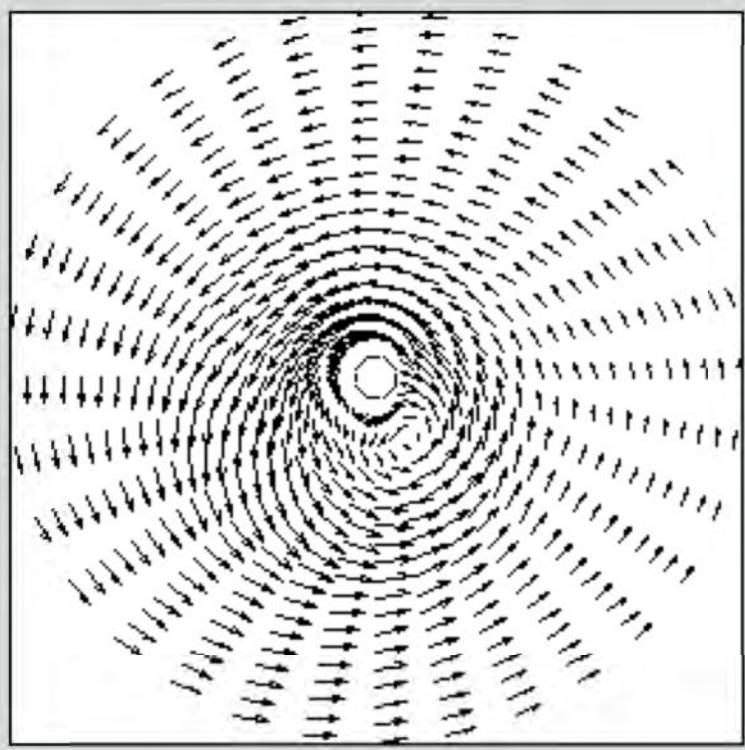


$\text{Ra} = 35000$

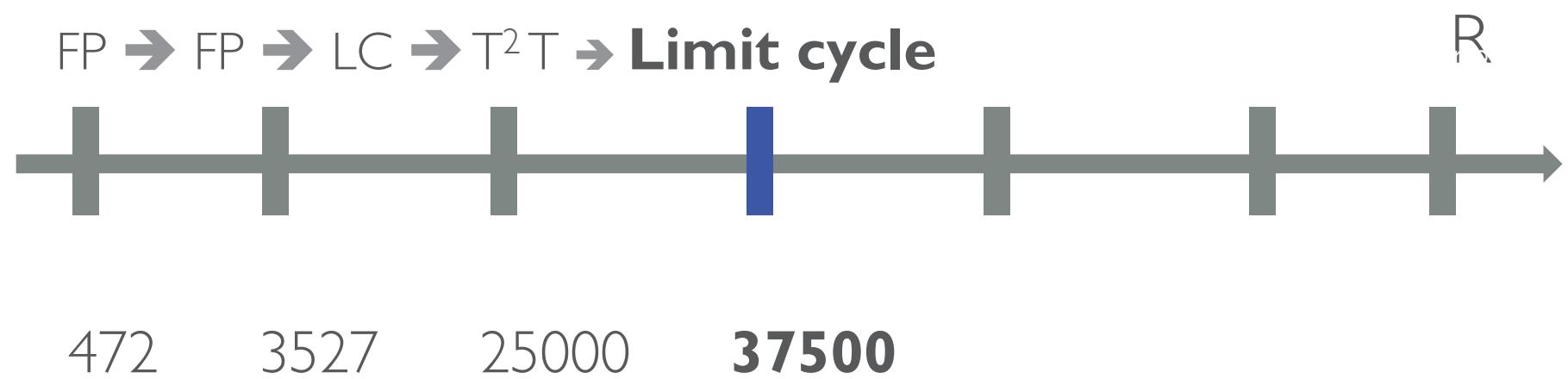


$\text{Ra} = 35000$

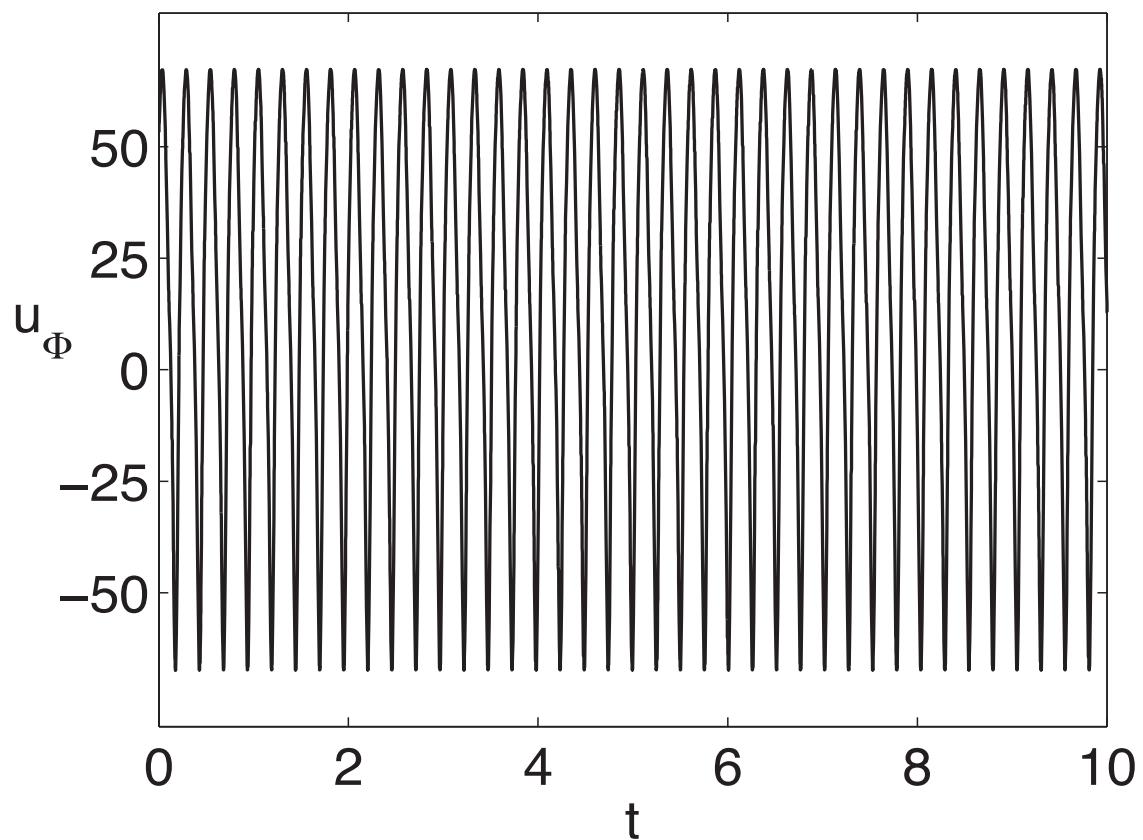




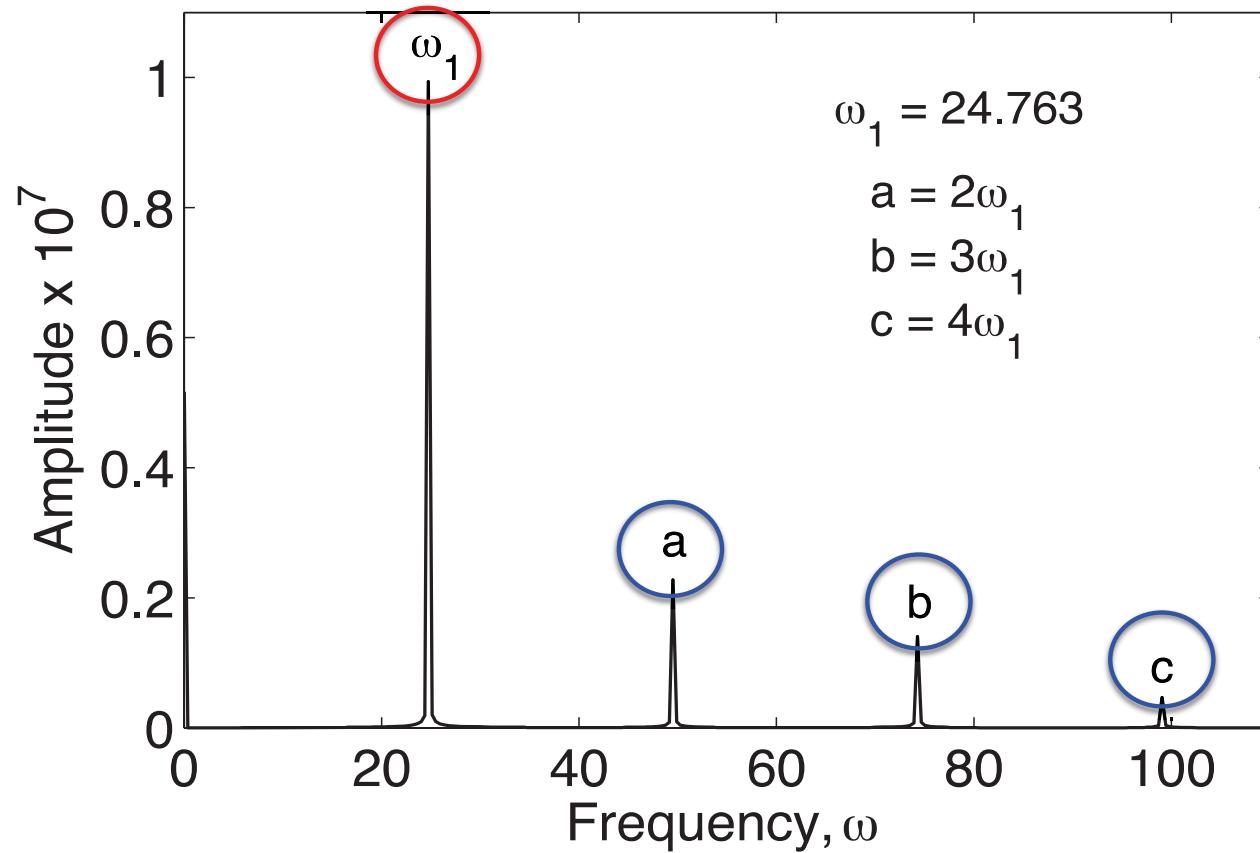
# Bifurcation diagram scheme



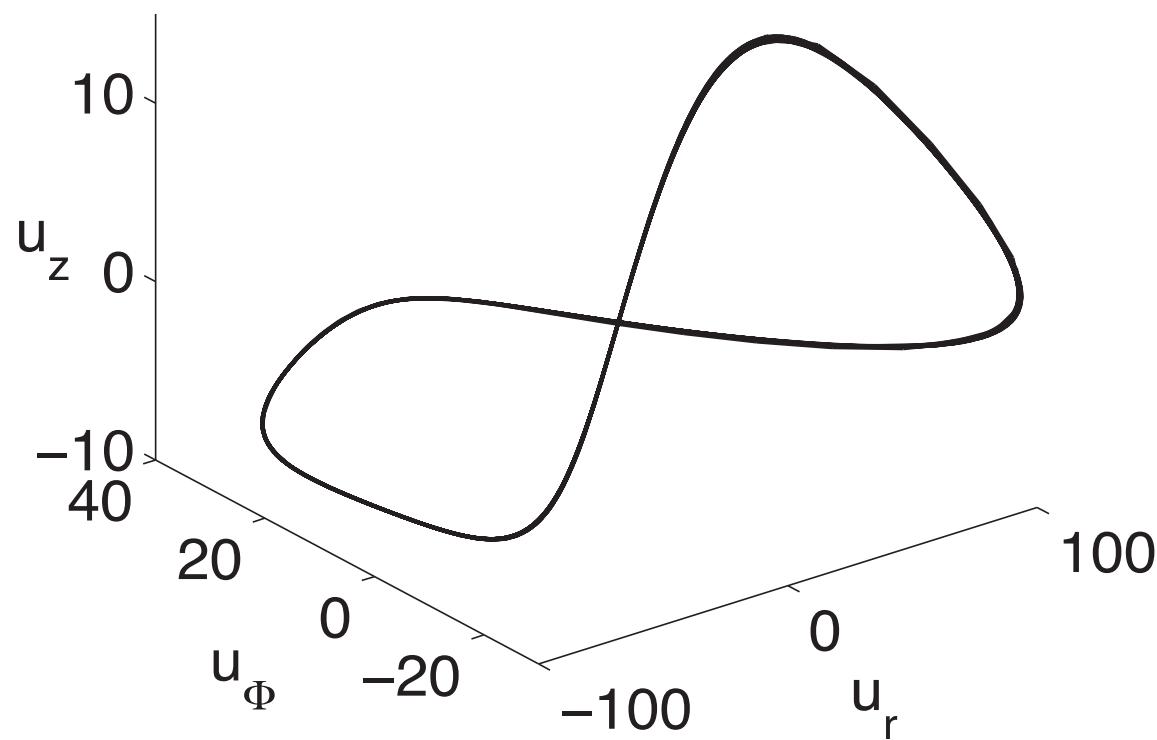
$\text{Ra} = 40000$



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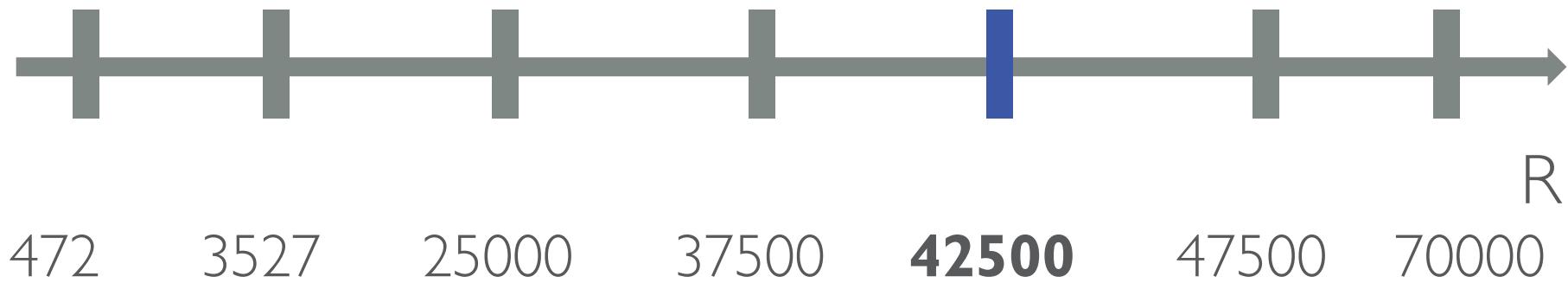


$\text{Ra} = 40000$

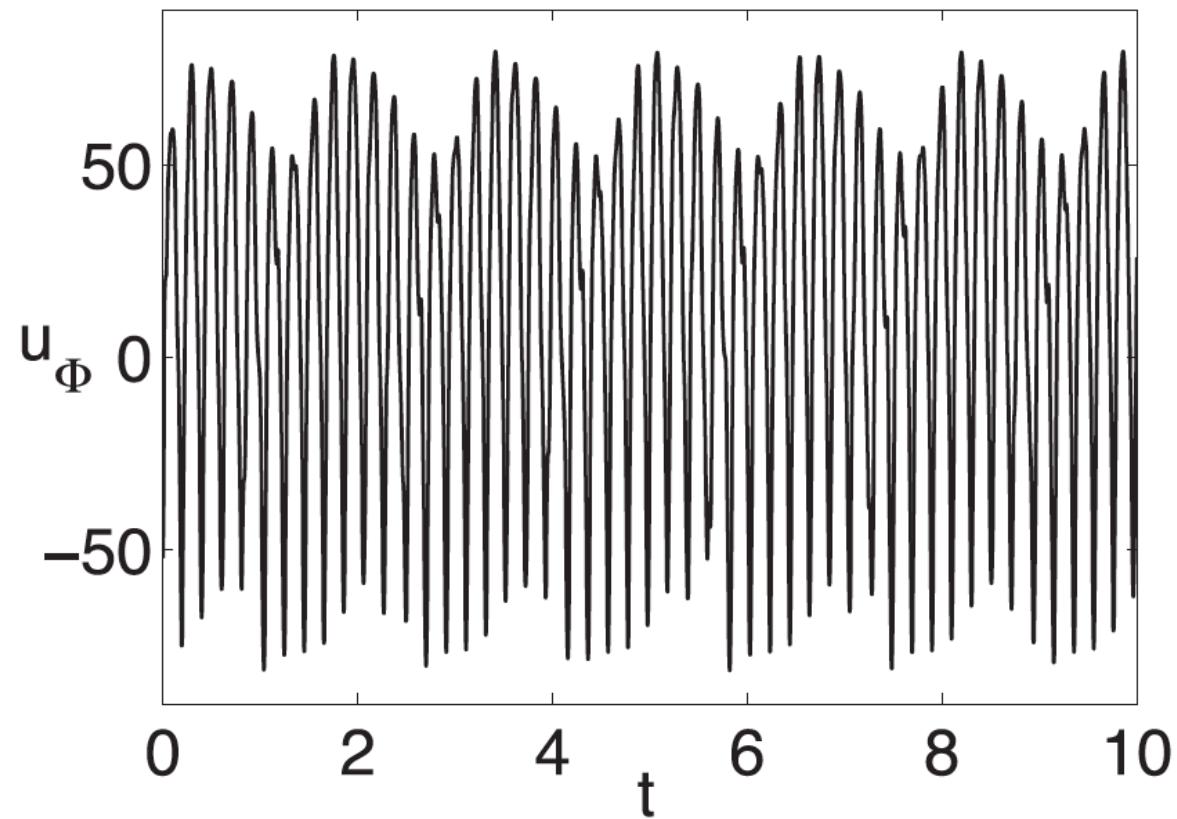


# Bifurcation diagram scheme

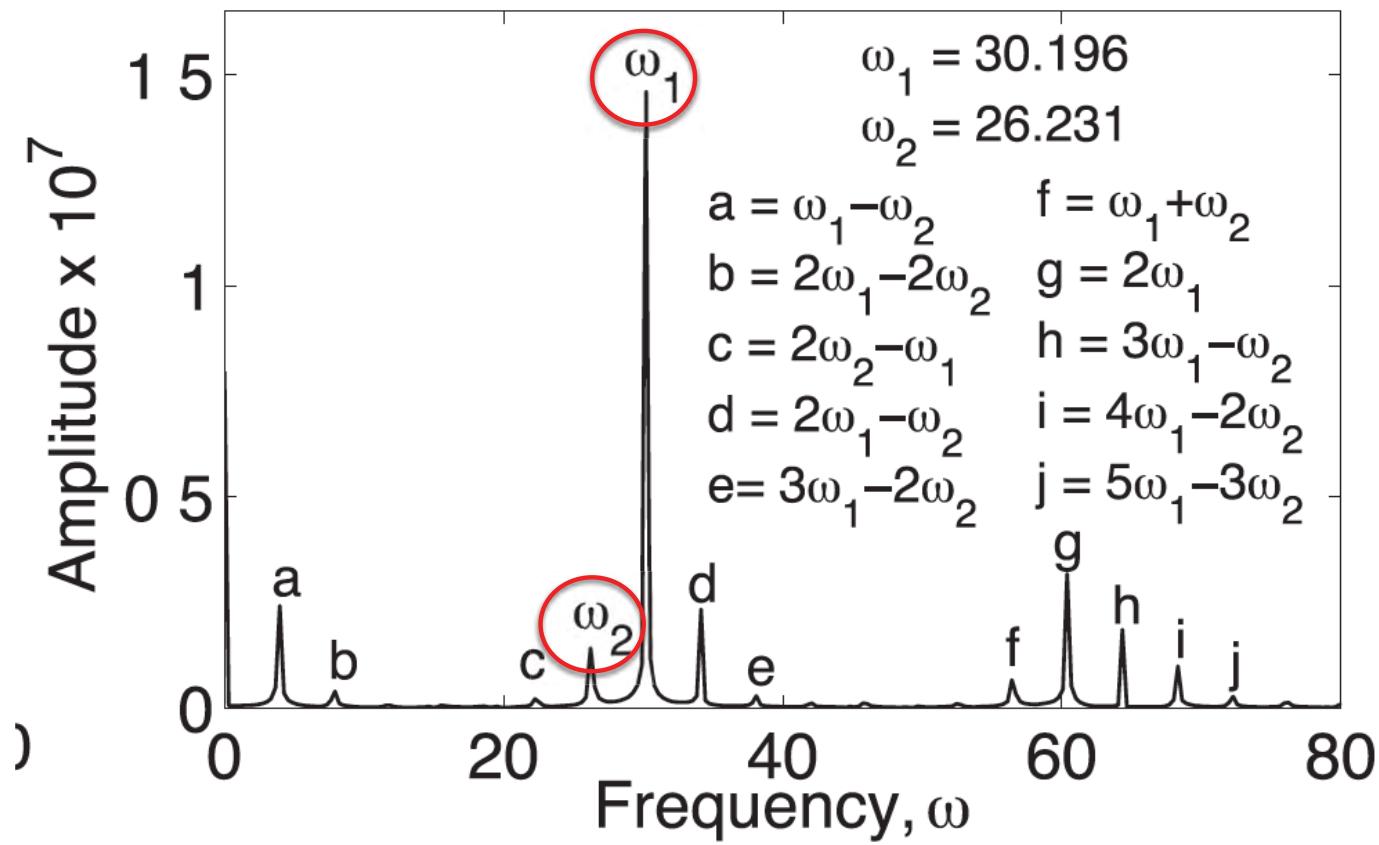
FP → FP → LC →  $T^2 T$  → LC →  **$T^2$  Torus**



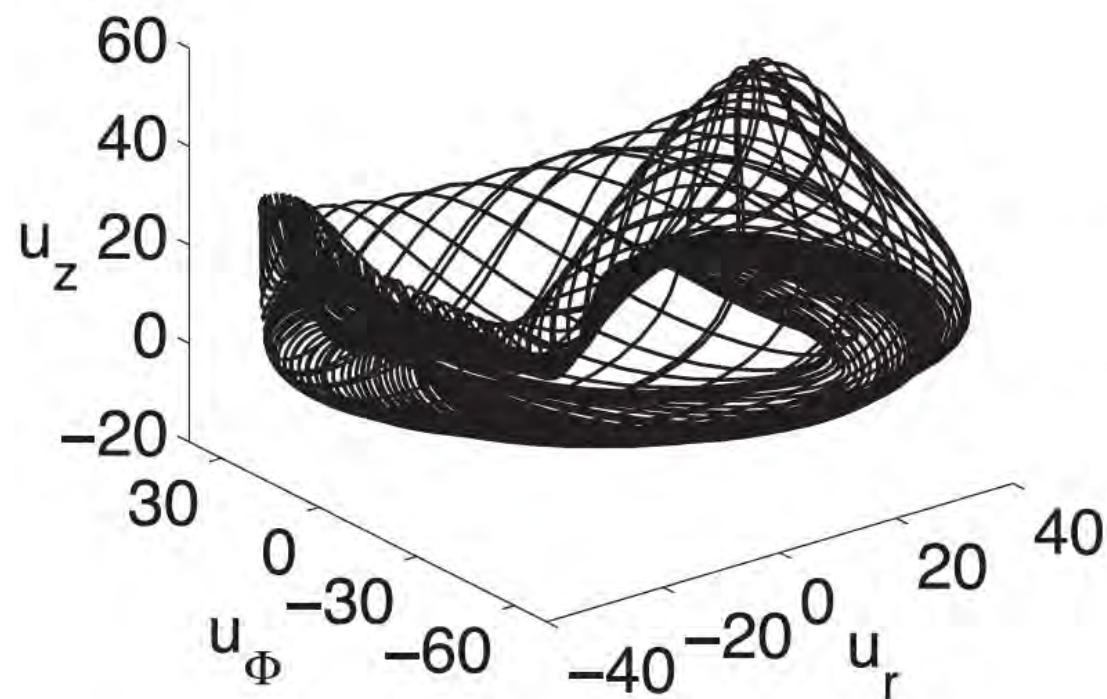
$\text{Ra} = 45000$



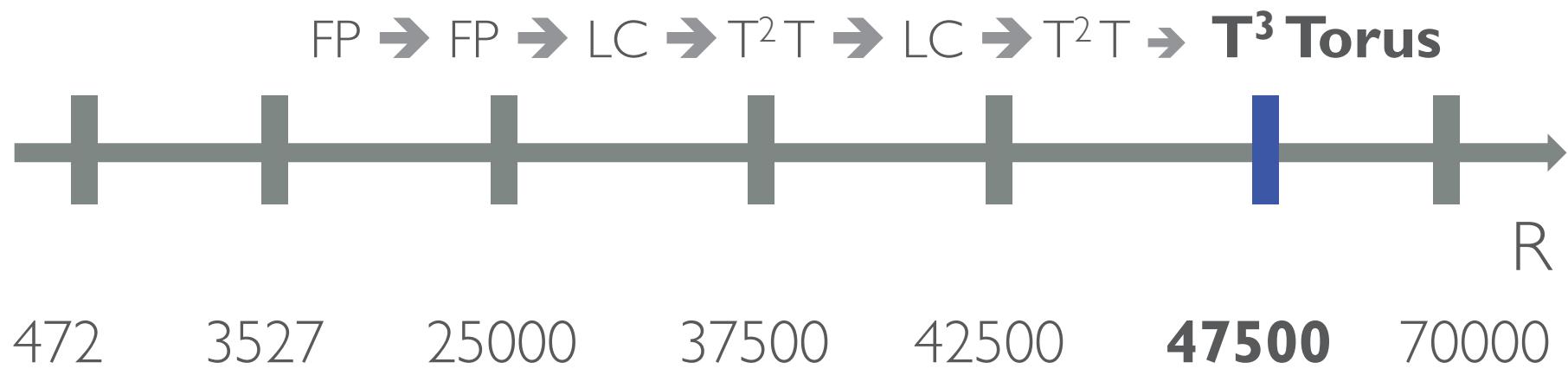
$\text{Ra} = 45000$



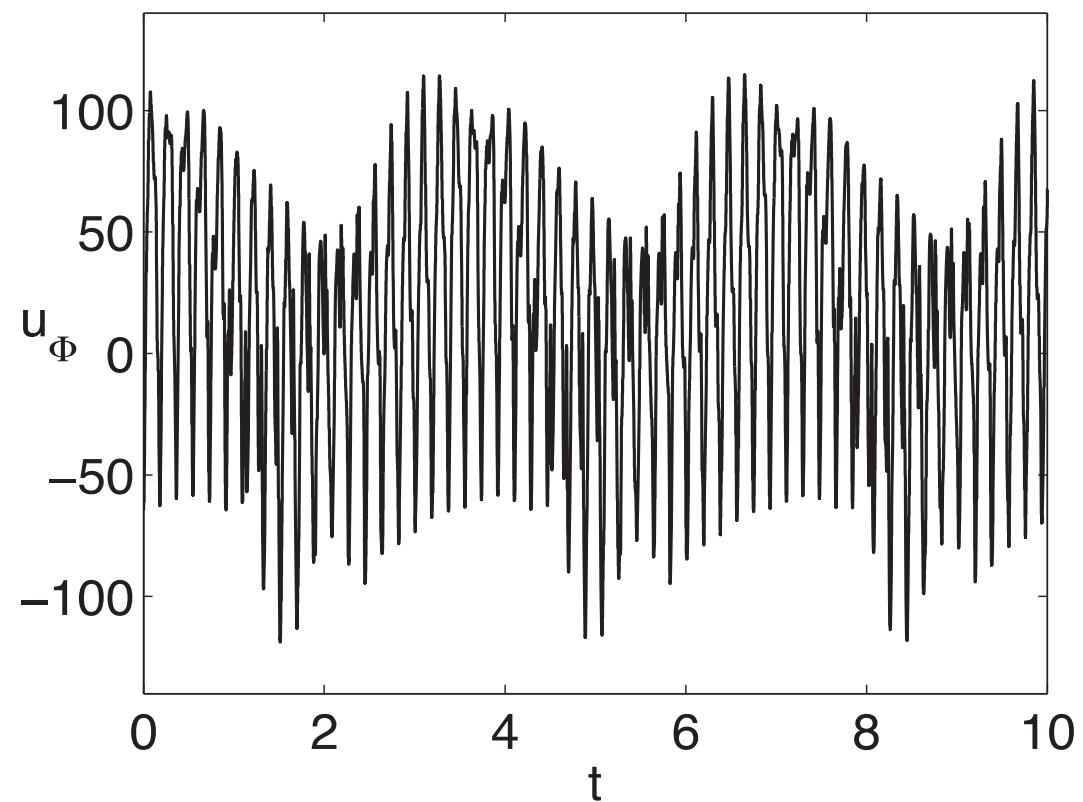
$\text{Ra} = 45000$



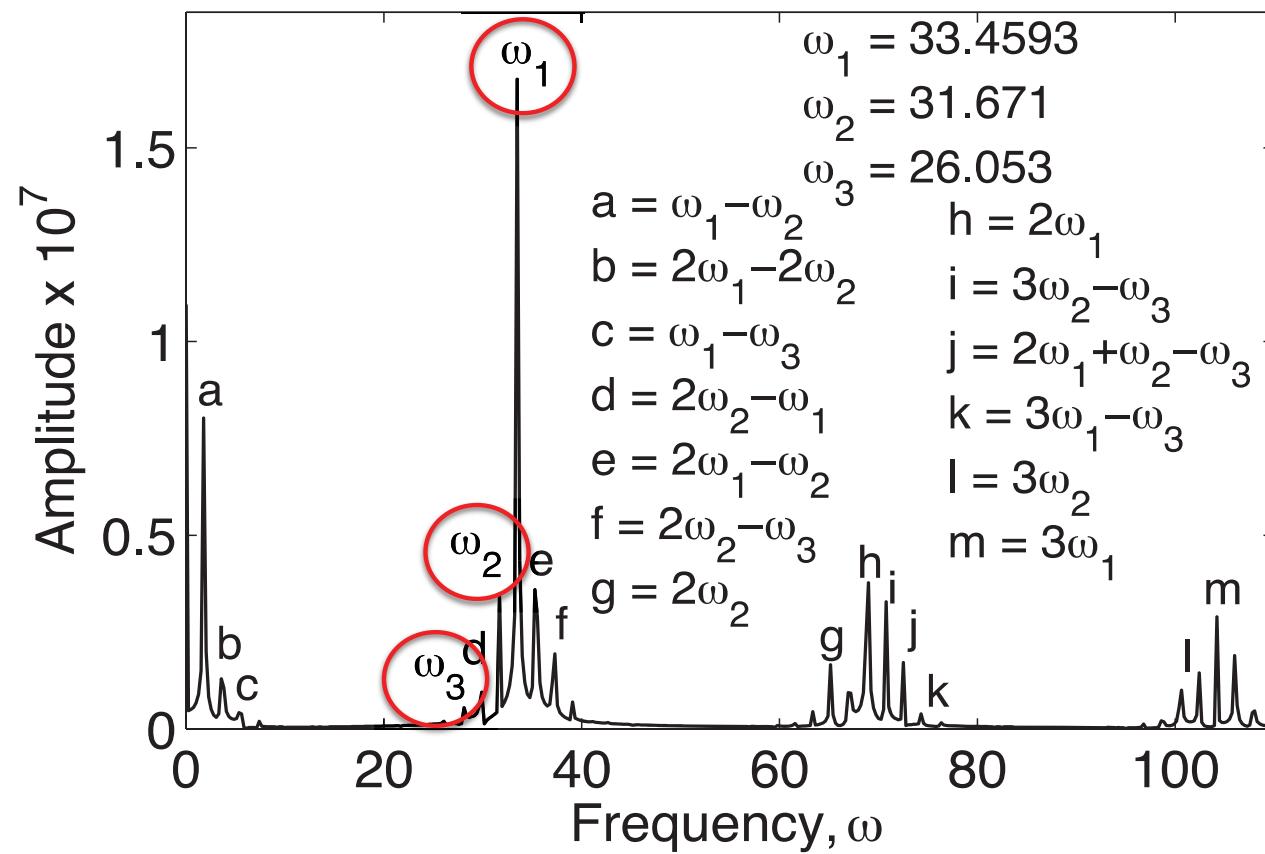
## Bifurcation diagram scheme



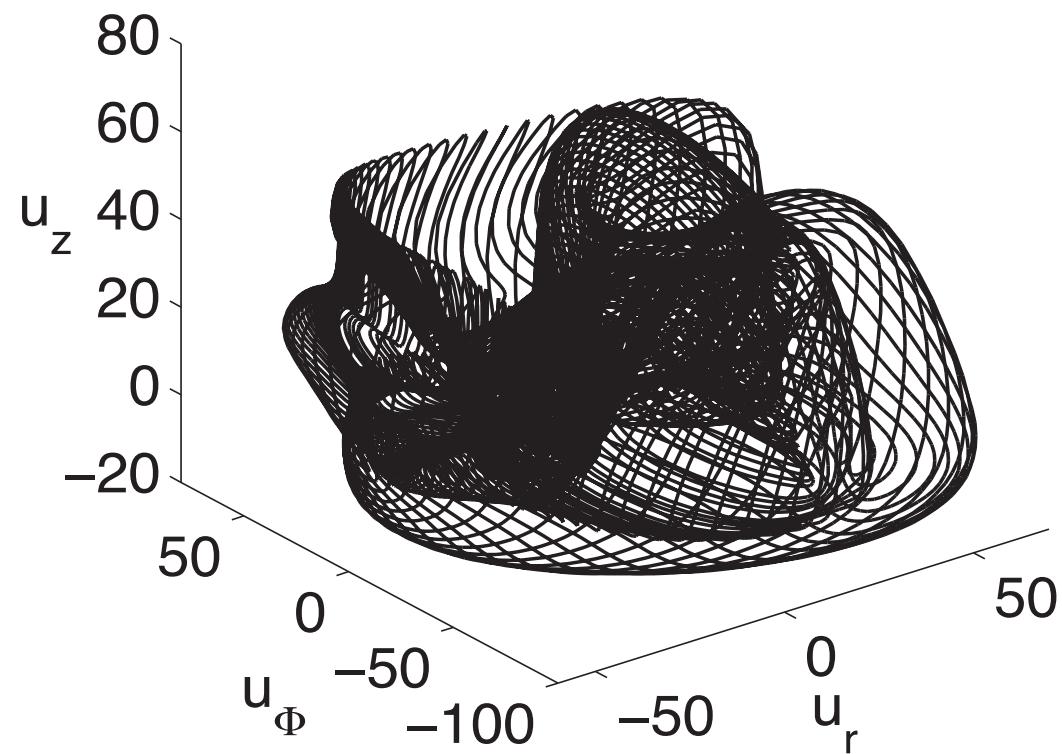
$\text{Ra} = 60000$

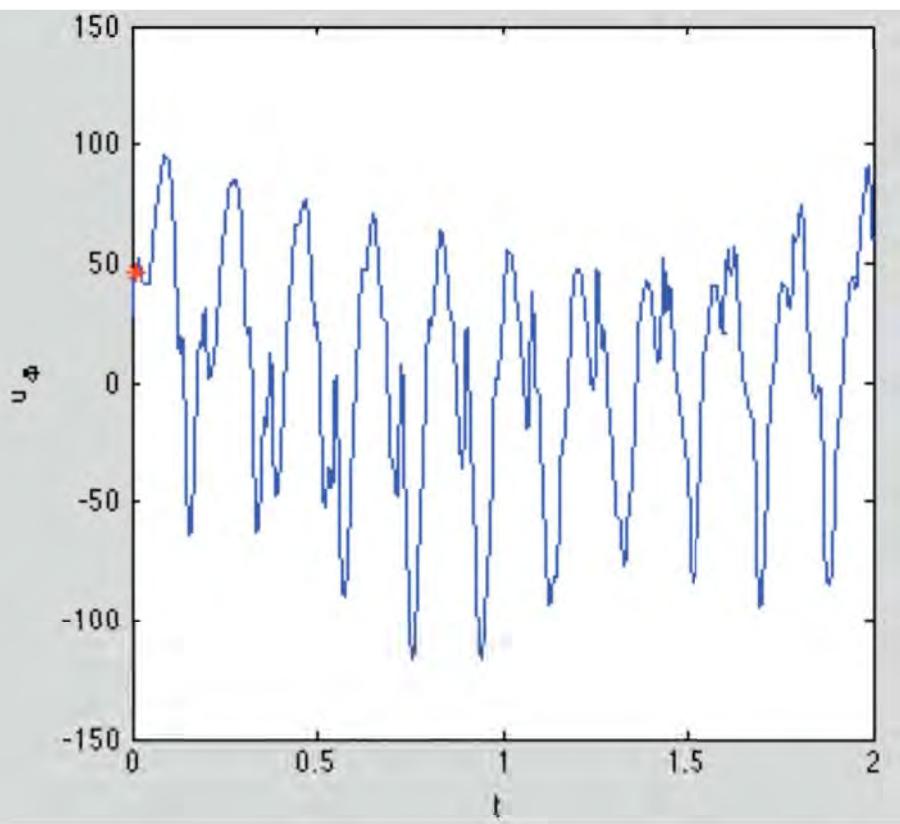
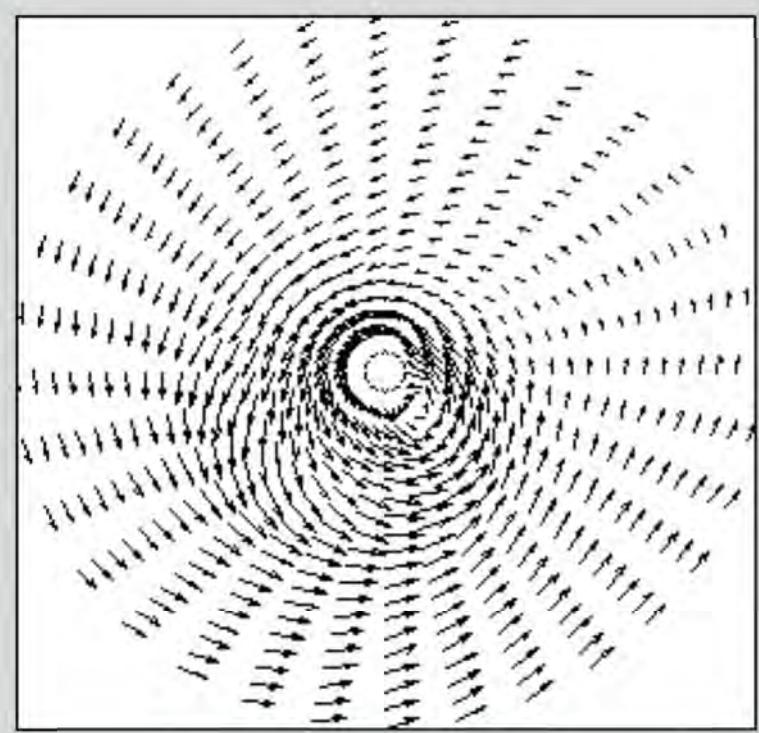


$\text{Ra} = 60000$



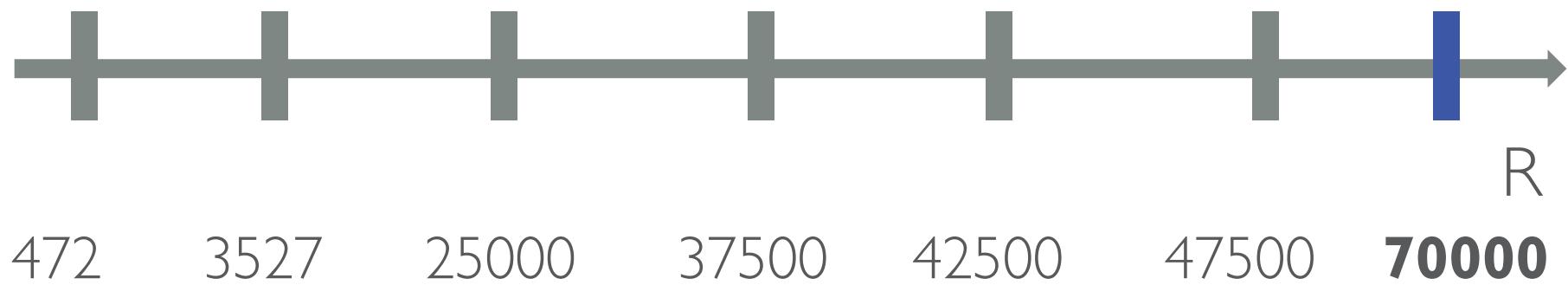
$\text{Ra} = 60000$





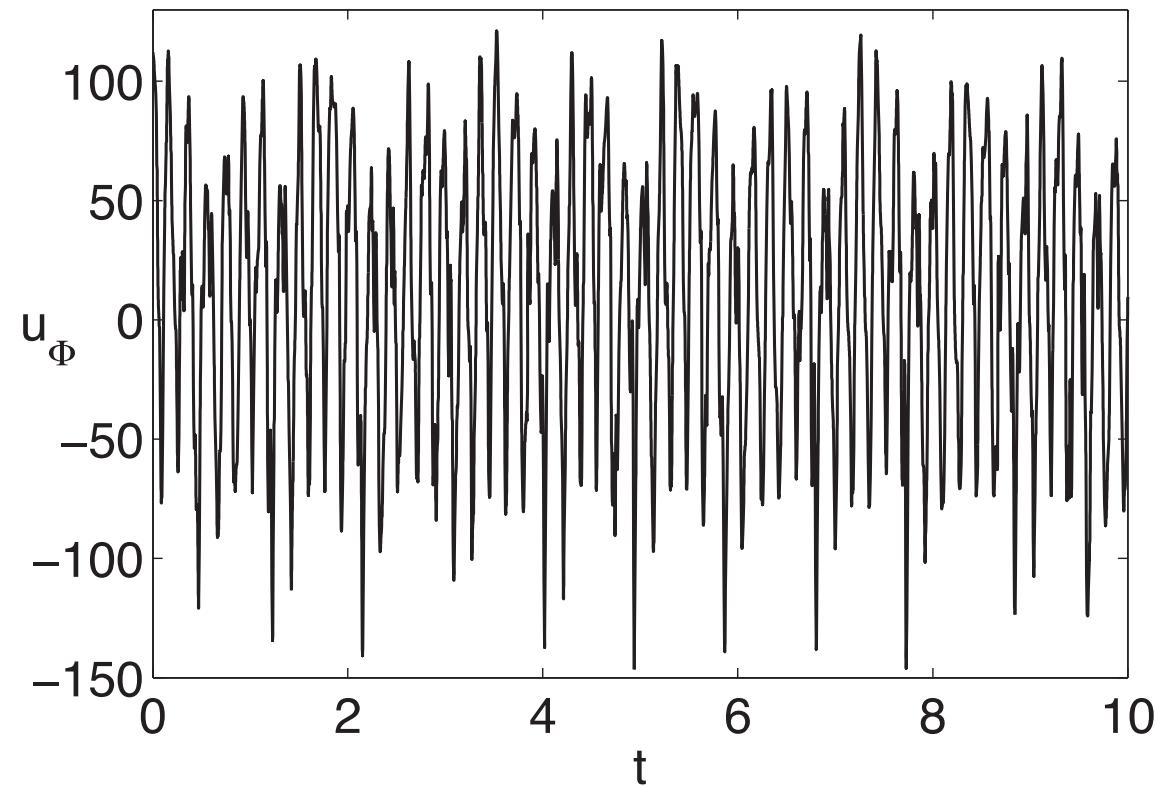
## Bifurcation diagram scheme

FP → FP → LC →  $T^2T$  → LC →  $T^2T$  →  $T^3T$  → **Strange attractor**

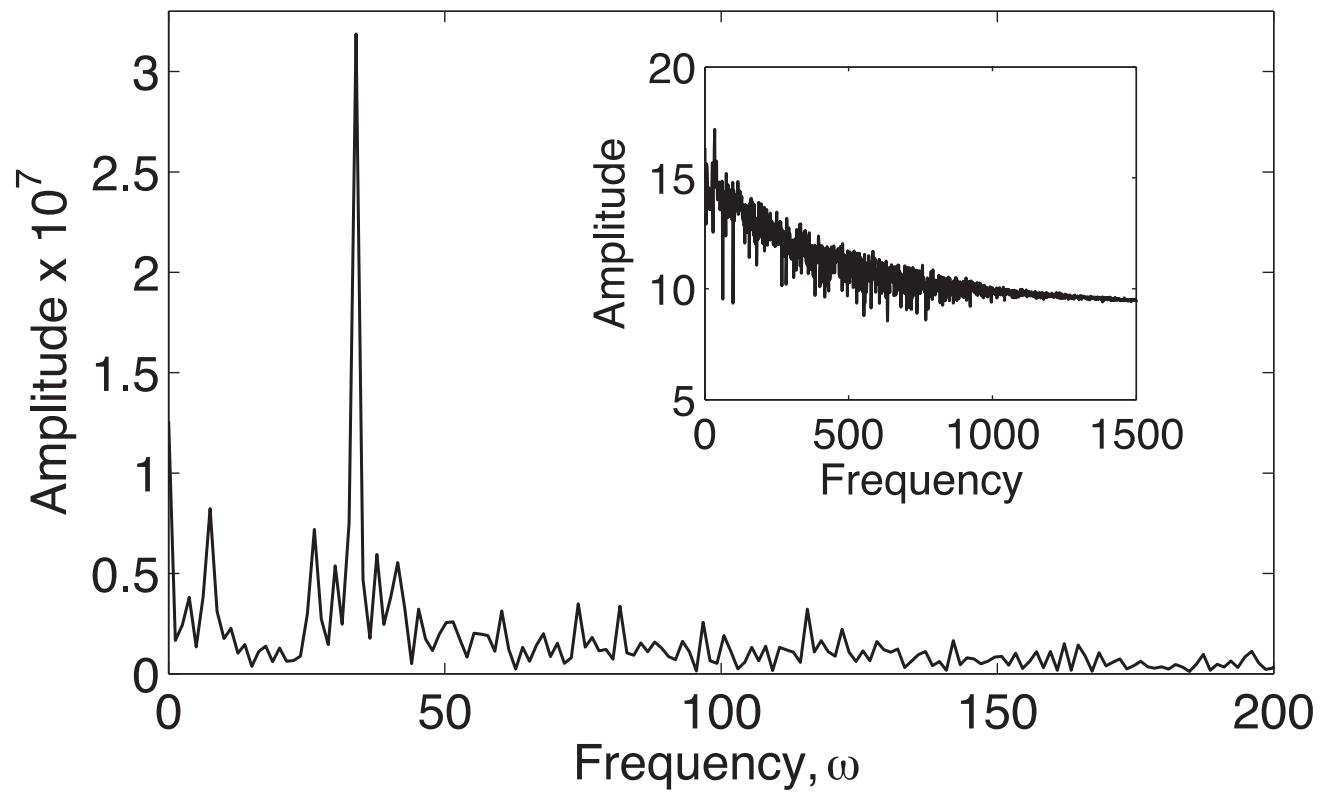


**Route to chaos → Ruelle-Takens-Newhouse scenario**

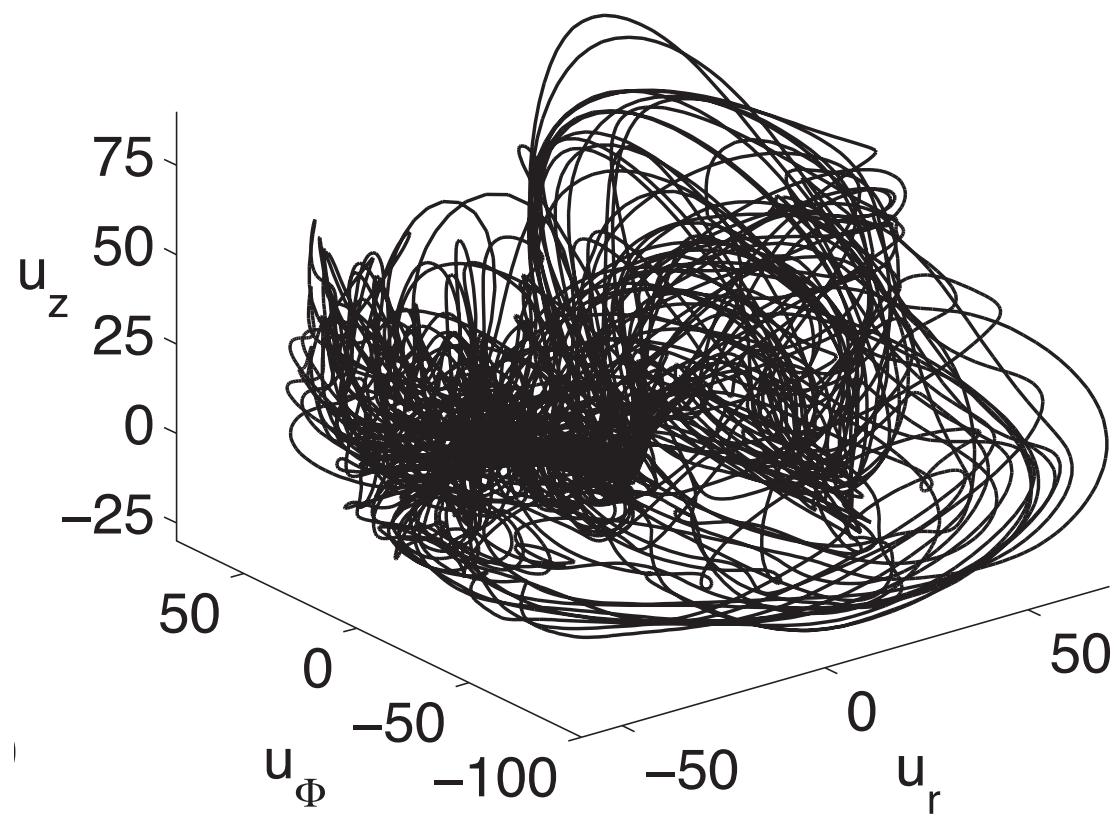
$\text{Ra} = 70000$



$\text{Ra} = 70000$



$\text{Ra} = 70000$

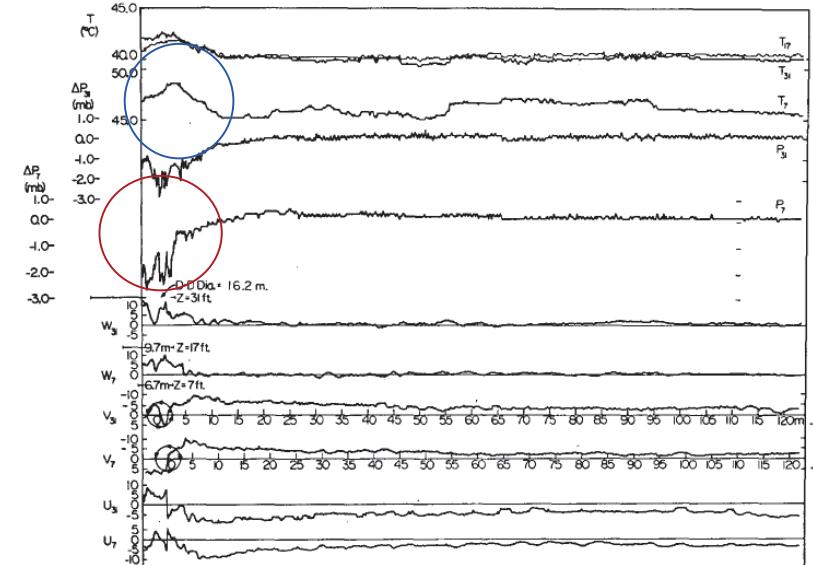
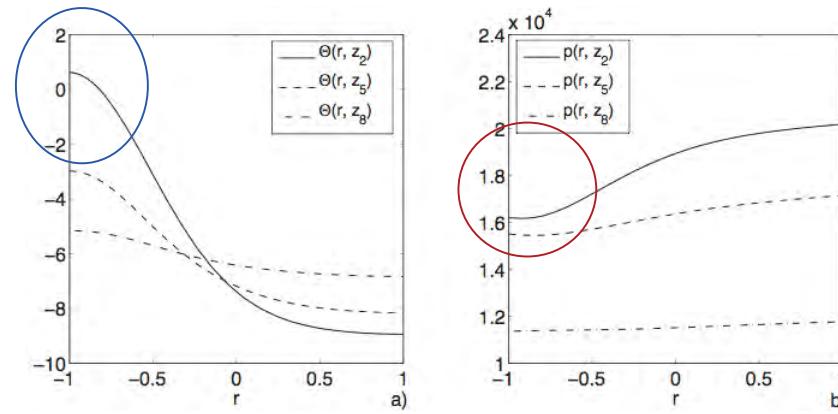


## **Comparison observations dust devils**

- Pressure, temperature and velocity fields
- Tilting of the axis of rotation
- Radius of maximum wind contraction
- Secondary whirls

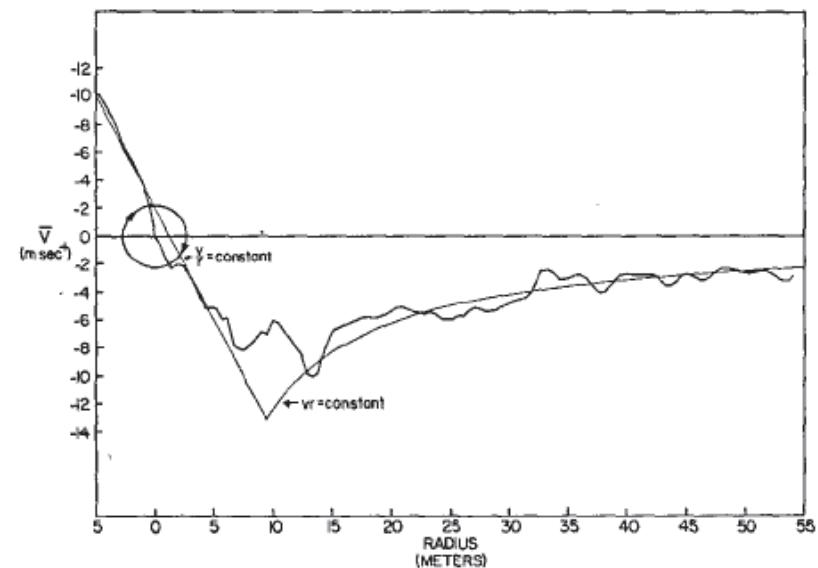
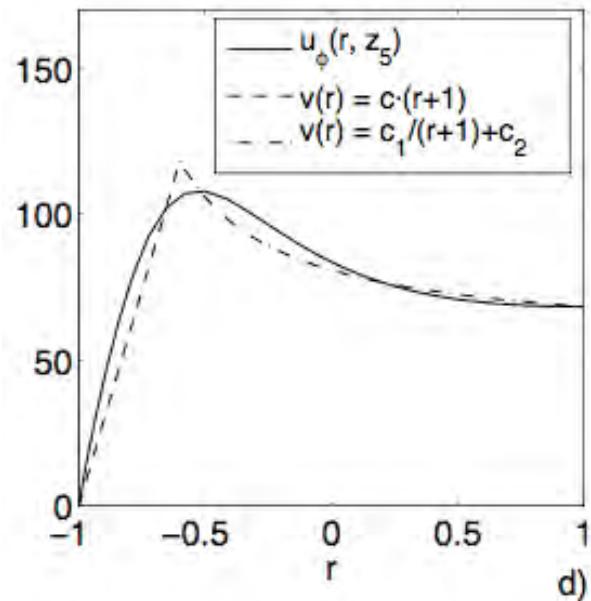
# Qualitative comparison with dust devils: pressure and temperature

- A low pressure region in the center
- Warm core



# Qualitative comparison with dust devils: tangential velocity

Rankine structure



# Qualitative comparison with dust devils

$$\Delta T_h \uparrow$$



$$\Delta p \uparrow$$



$$u_\phi \uparrow$$

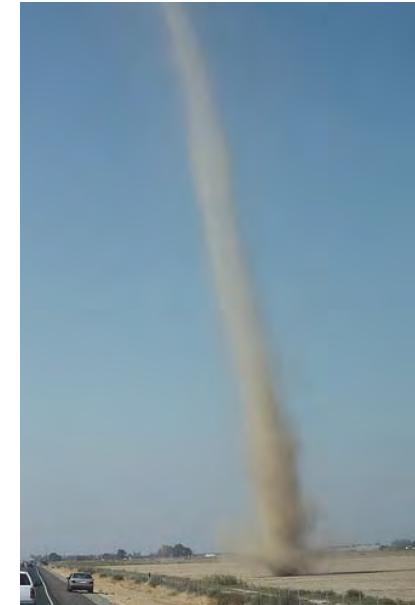
$$\Delta T_h \downarrow$$

(cooling at the core)

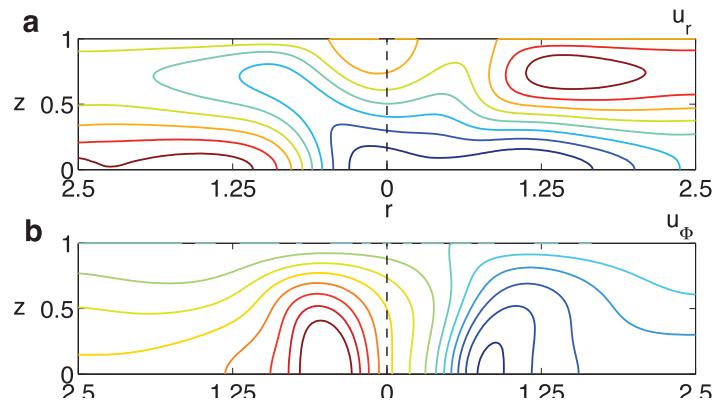


$$u_\phi \downarrow$$

(the dust devil disappears)

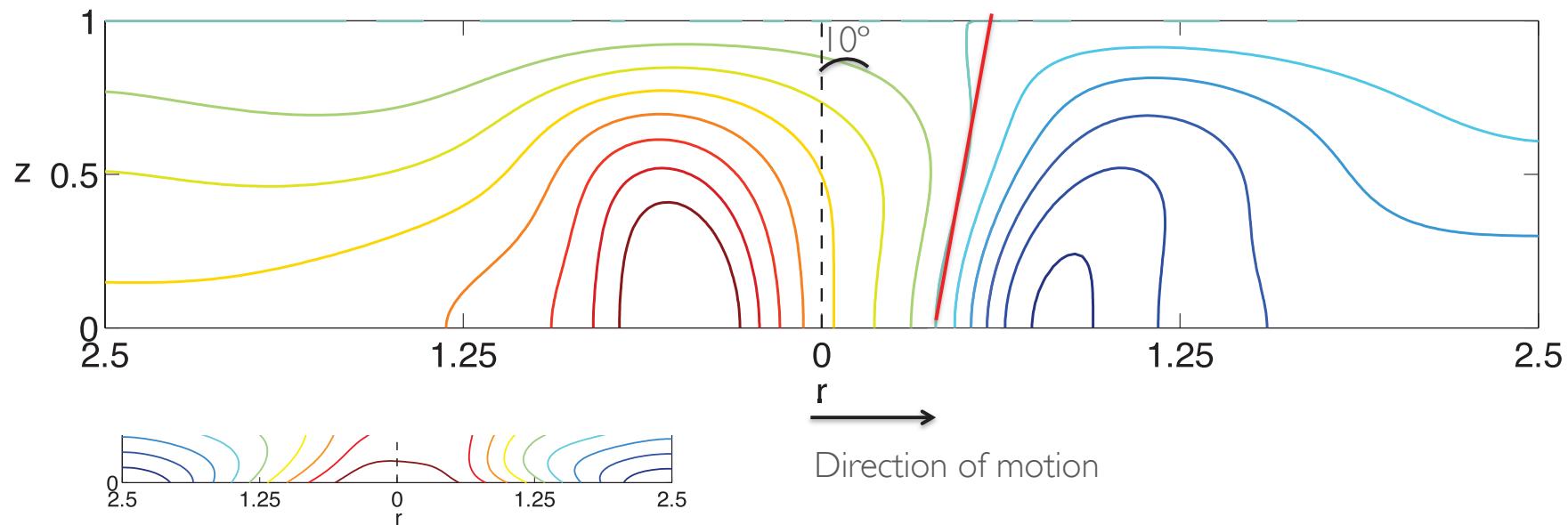


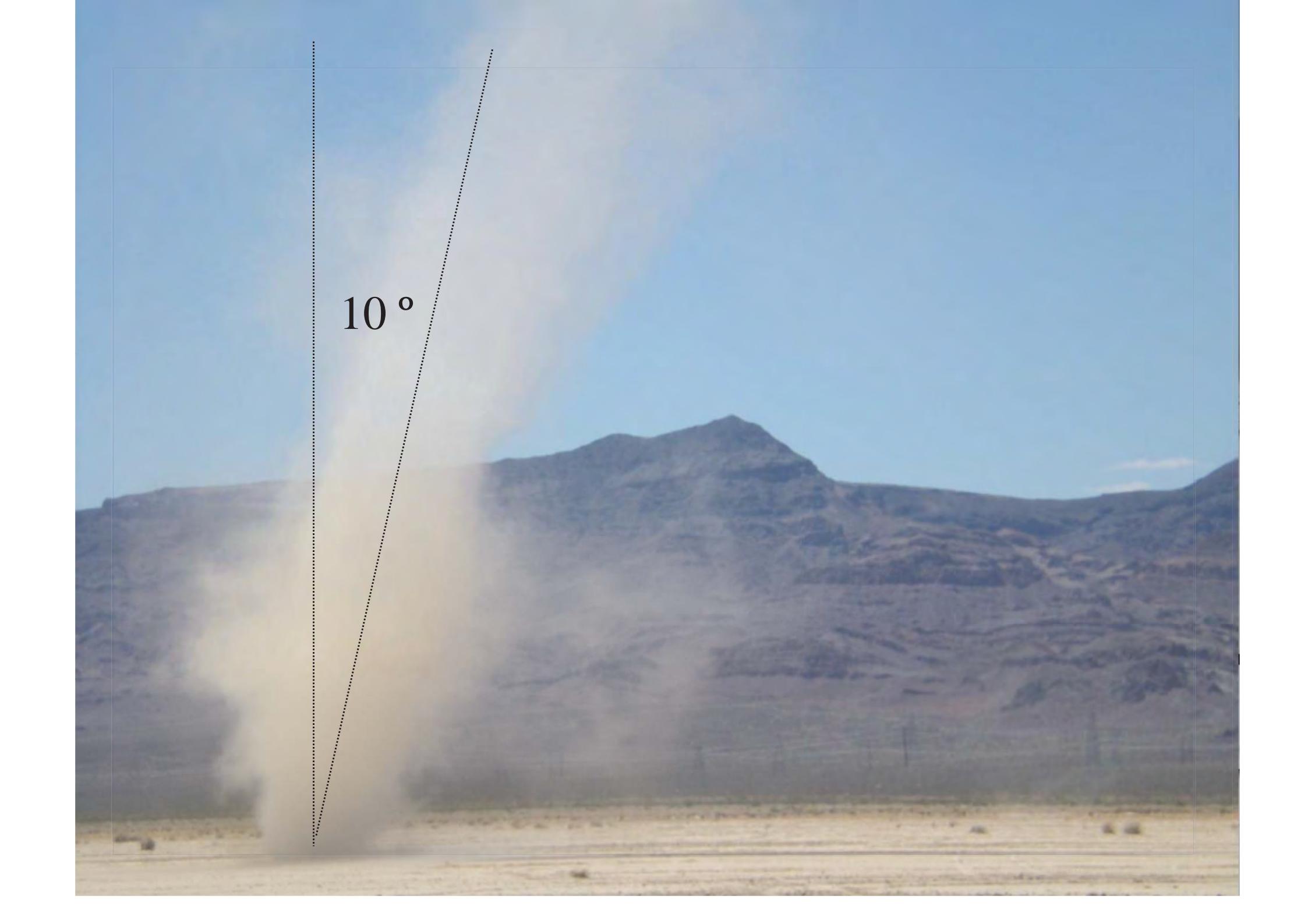
# Tilting



Inclination  $10^\circ$

-McGinnigle, Weather, 1966  
-Mattson, Weather, 1993

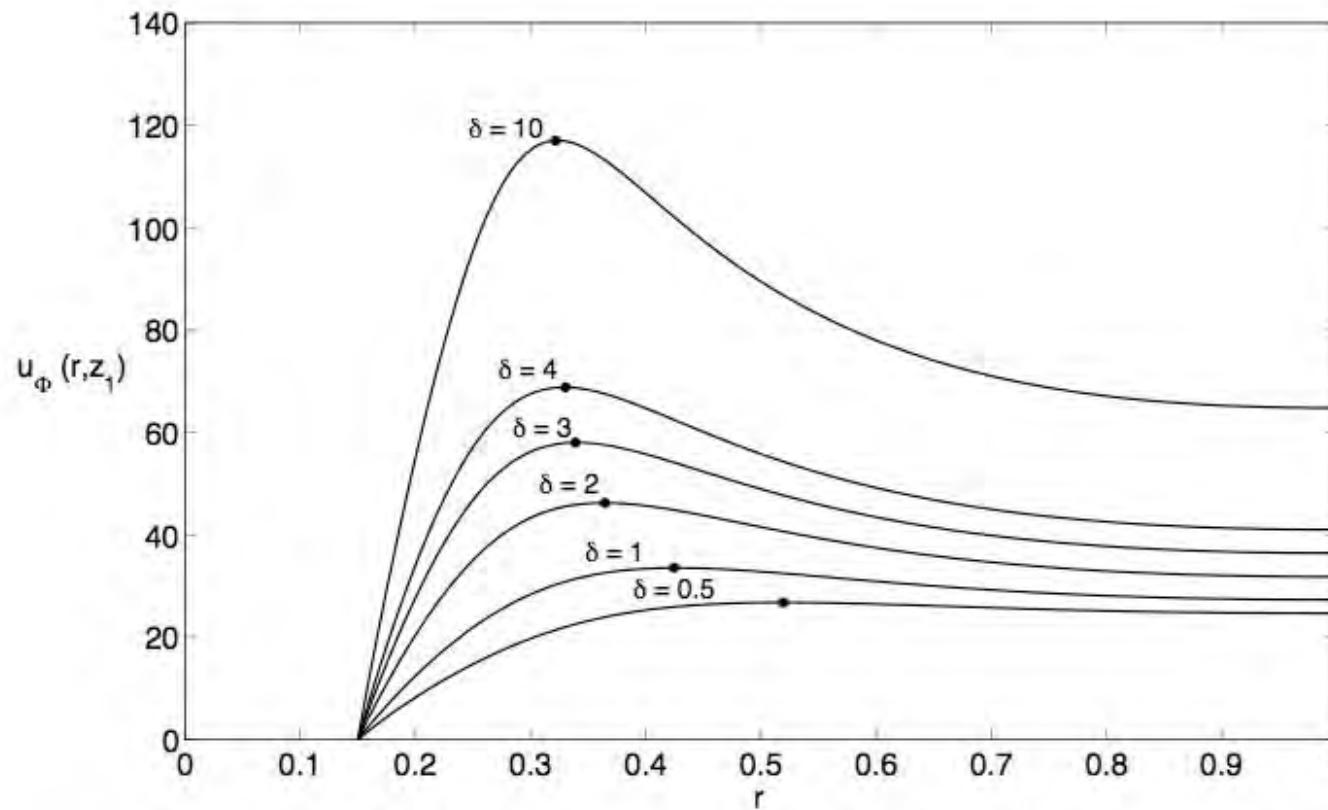




$10^{\circ}$



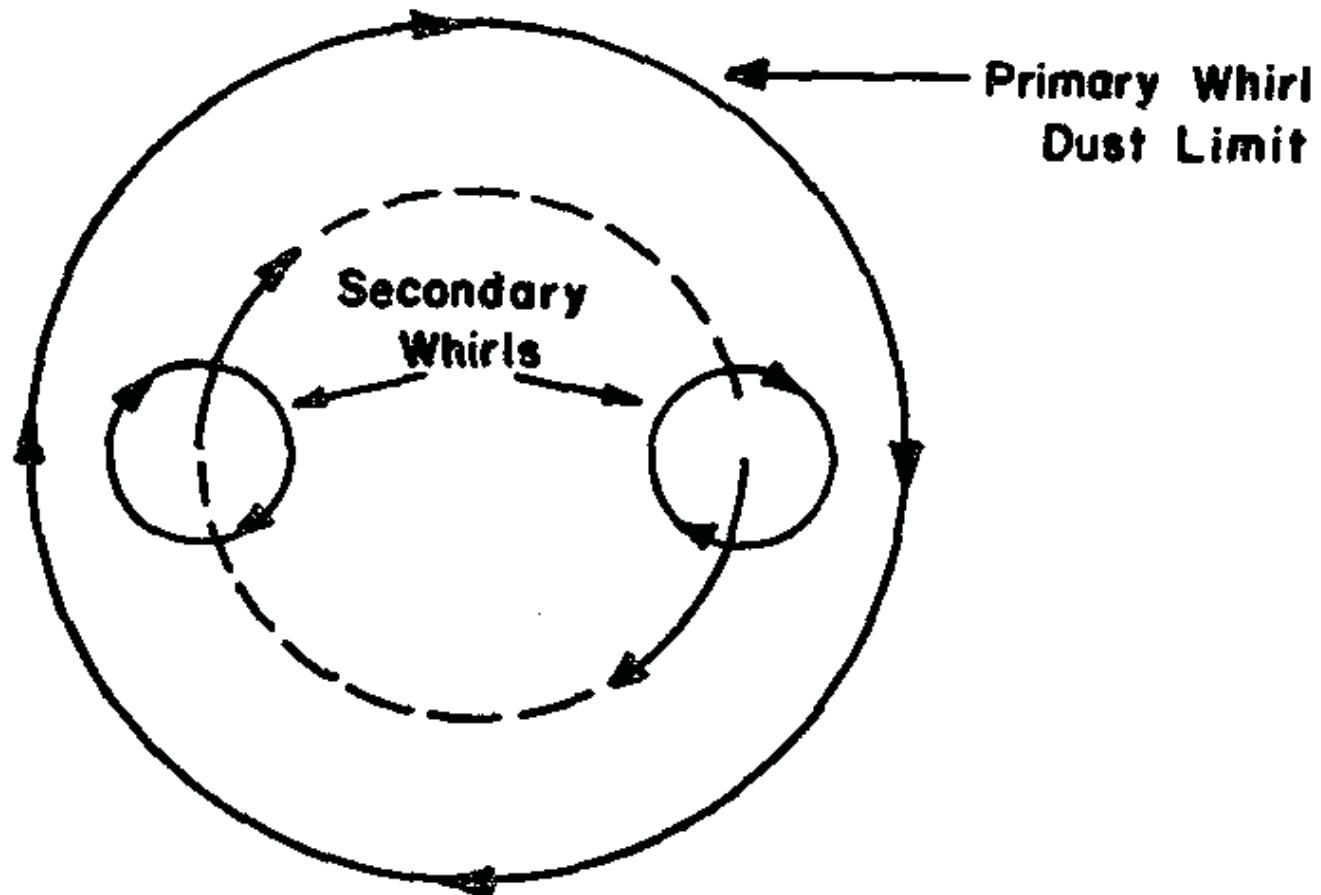
# RMW contraction

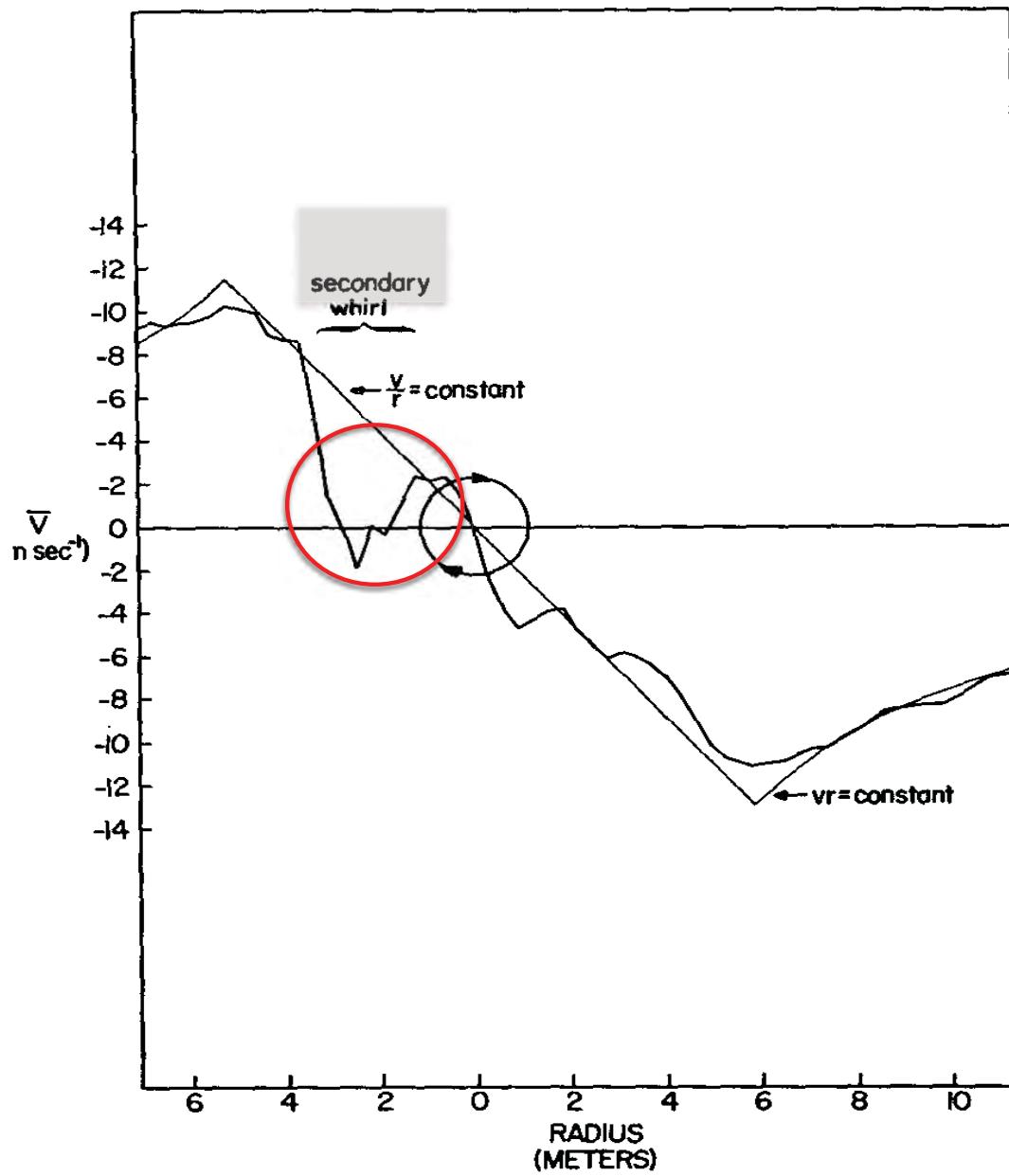


**Fig. 6.** (a) Critical Rayleigh number  $R_c$  and critical wavenumber  $k_c$  as a function of  $\delta$  for  $\Gamma = 0.5$ ; (b) critical Rayleigh number  $R_c$  and critical wavenumber  $k_c$  as a function of  $\Gamma$  for  $\delta = 1$ . Empty circles correspond to real eigenvalues while filled ones stand for complex eigenvalues.

# **Secondary whirls**

Peter C. Sinclair, Journal of the Atmospheric Sciences 30, 1599-1619 (1973)





(a)

# Conclusions

Vortical structures (spiral up motion) can be generated by a thermoconvective instability in a Rayleigh-Bénard problem (time dependent 3D numerical results)

- in a cylinder and in a cylindrical annulus
- heated non-homogeneously from below
- open outer boundary

Similar to dust devils and cyclones

- tilting
- RMW
- effect of thermal gradients
- secondary whirls

# Third International Workshop NONLINEAR PROCESSES IN OCEANIC AND ATMOSPHERIC FLOWS

Thank you for your attention

Sponsored by: MINECO (Spain) and RDF (EU)

